



PERFORMANCE VERIFICATION OF SOLAR-FLARE PREDICTION MODELS

from Climatology to Skill and from Forecast Probabilities to Certainty

Manolis K. Georgoulis
RCAAM of the Academy of Athens

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- * A-EFFort (ESA/SSA): a-effort.academyofathens.gr
- * FLARECAST (EC/H2020): flarecast.eu



SCIENCE FOR SPACE WEATHER

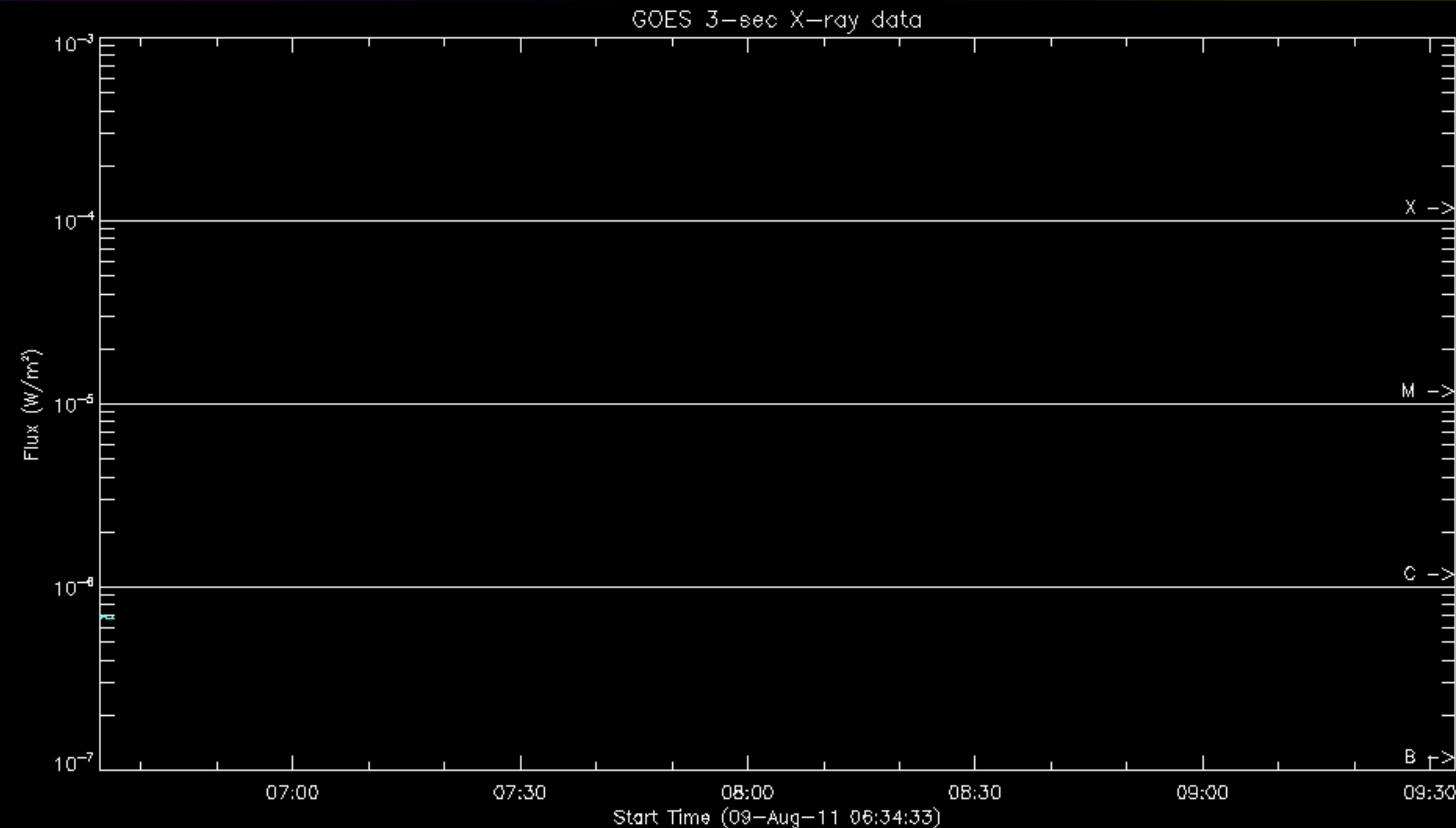
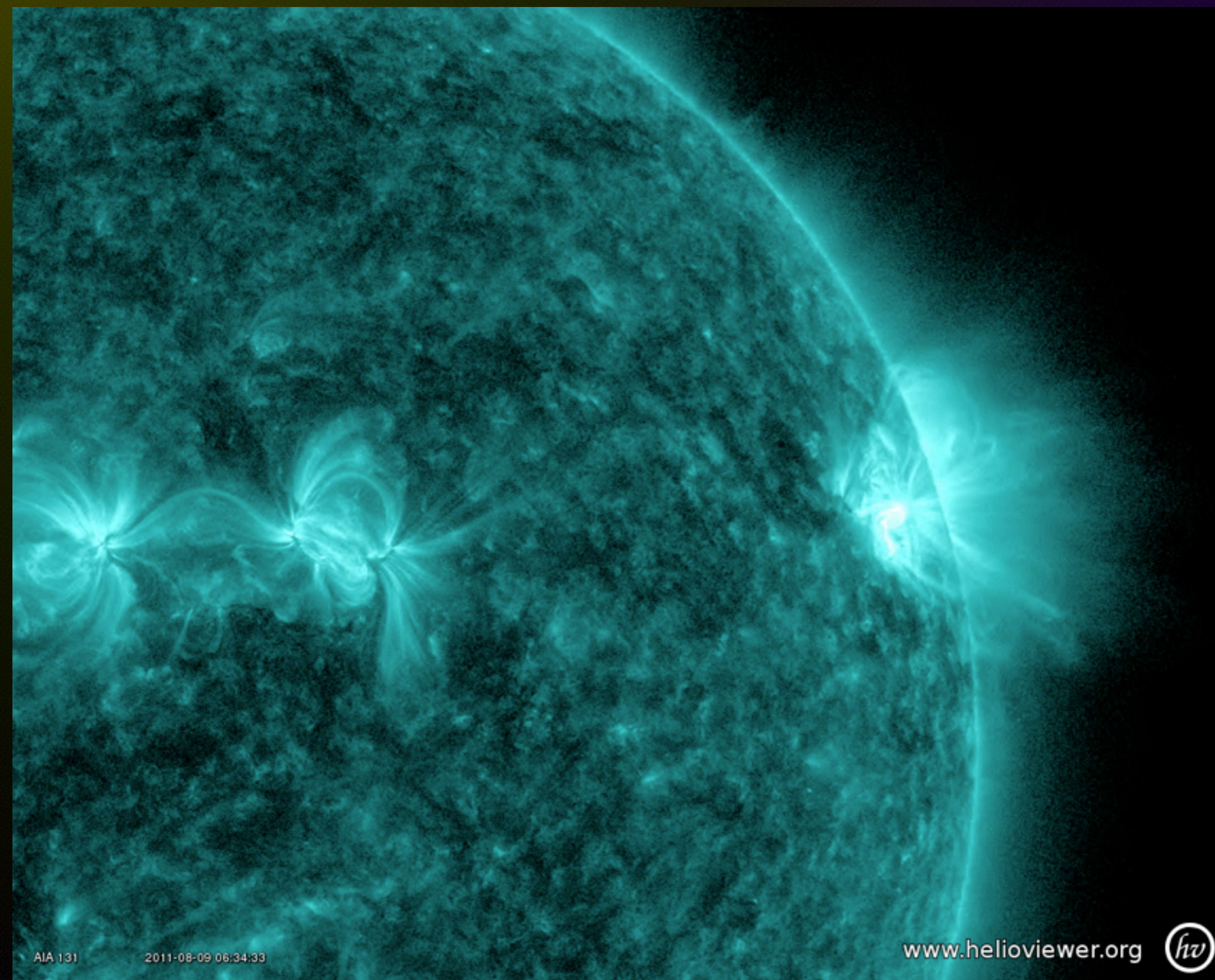
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OUTLINE

- ★ Solar flares and the prediction challenge
- ★ Validation of flare prediction methods
 - *Dichotomous validation*
 - *Probabilistic validation*
- ★ Tailoring prediction methods to the customers' needs
 - *Multi-variable forecasting*
 - *Ensemble forecasting*
- ★ Conclusion



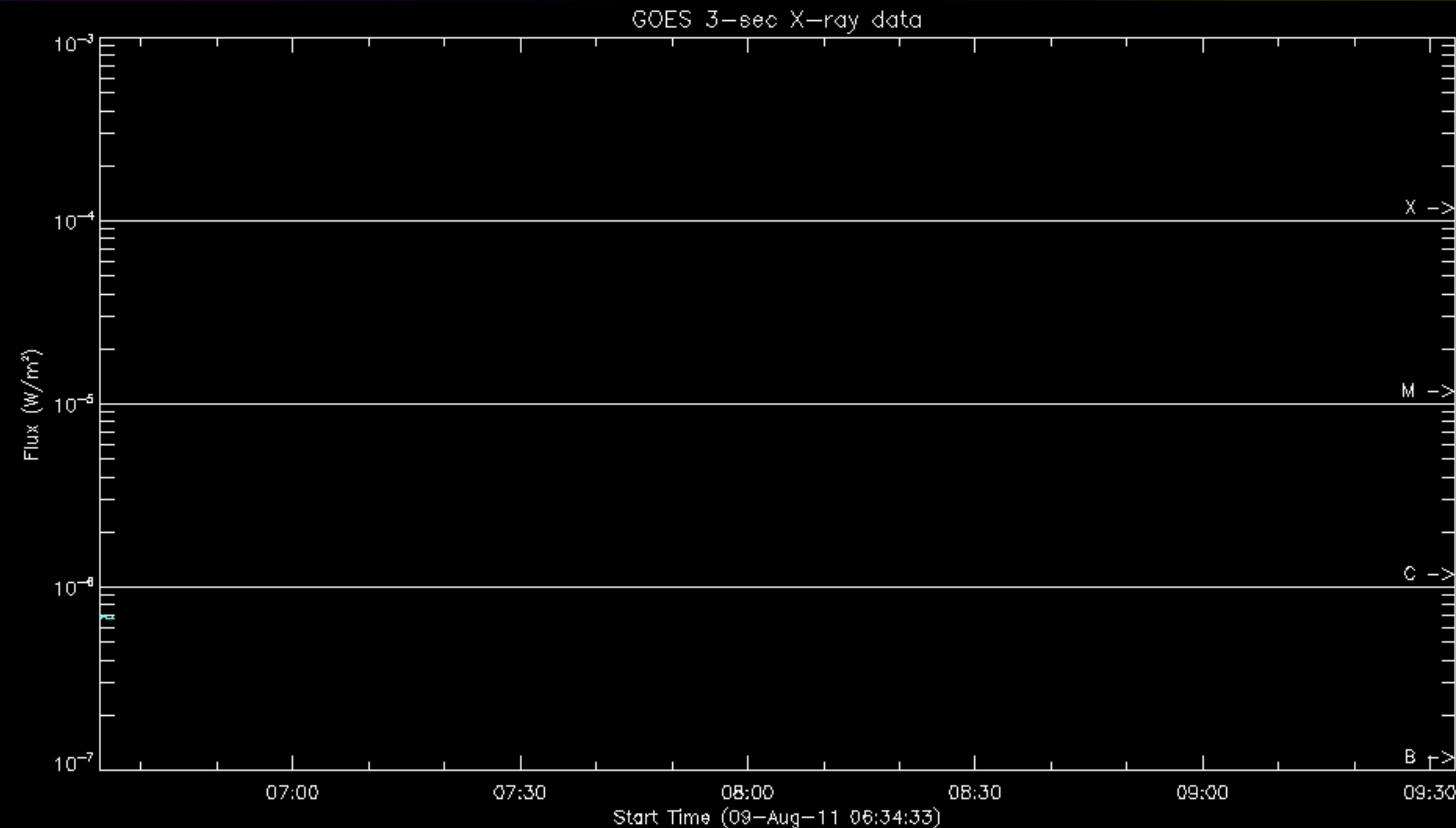
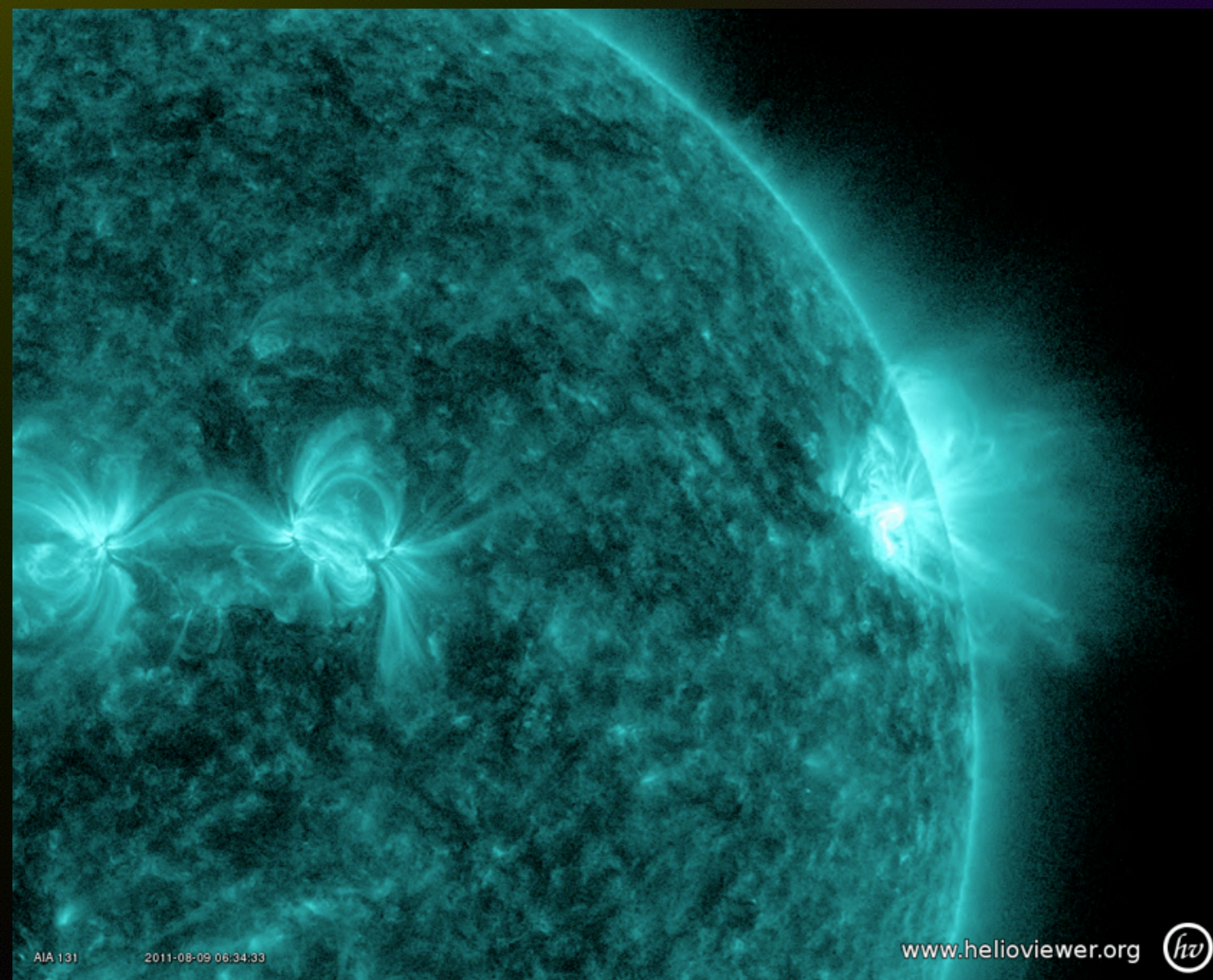
SOLAR FLARES: THE SINGLE ...



A sudden commencement of enhanced, localized electromagnetic emission extending over practically the entire range of the electromagnetic spectrum, from γ -ray to radio wavelengths



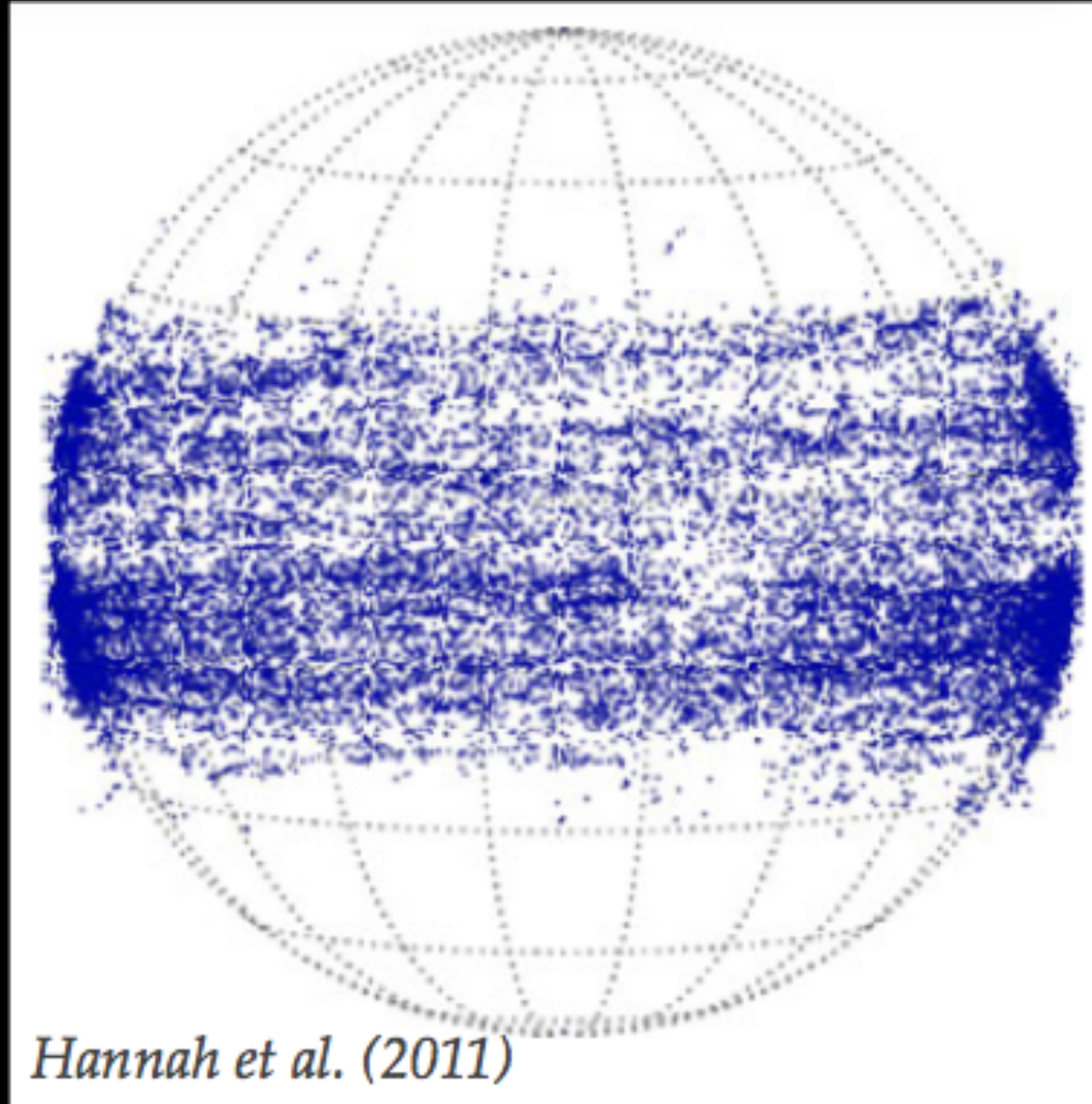
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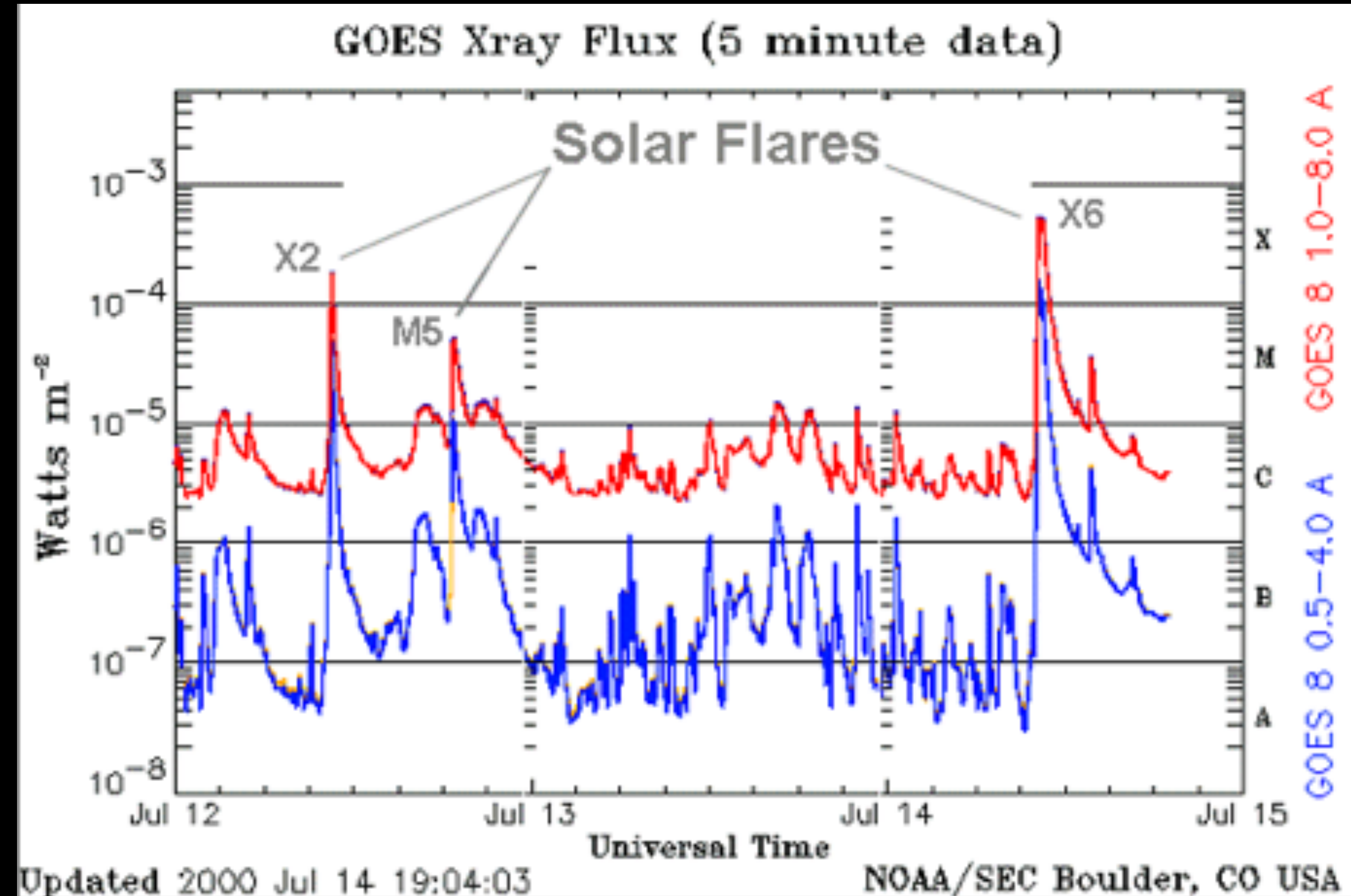
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SOLAR FLARES: ... AND THE PLENTY



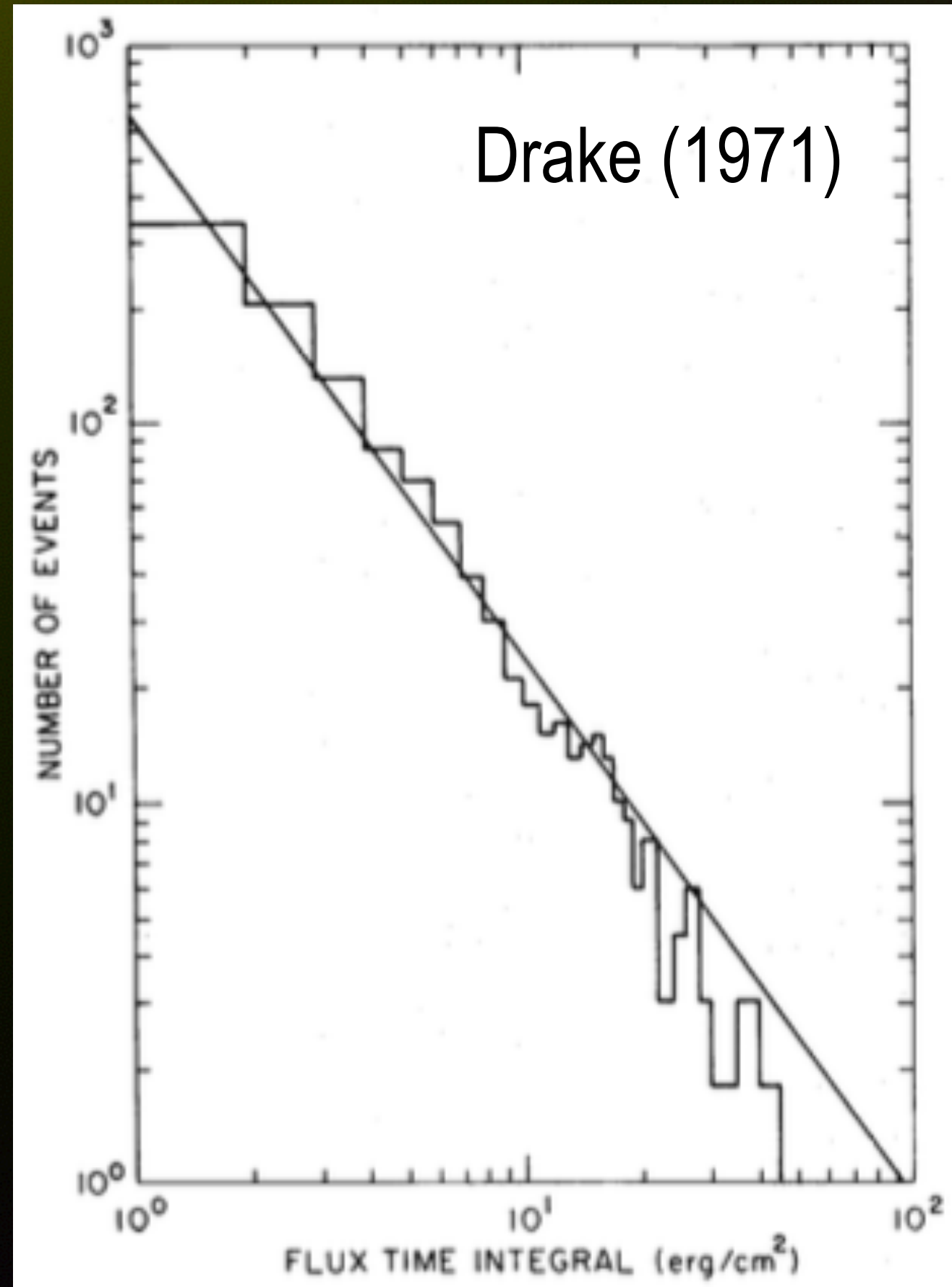
Location of some 27,000 flares from the RHESSI database



“Active” solar conditions over a 3-day period in July 2000



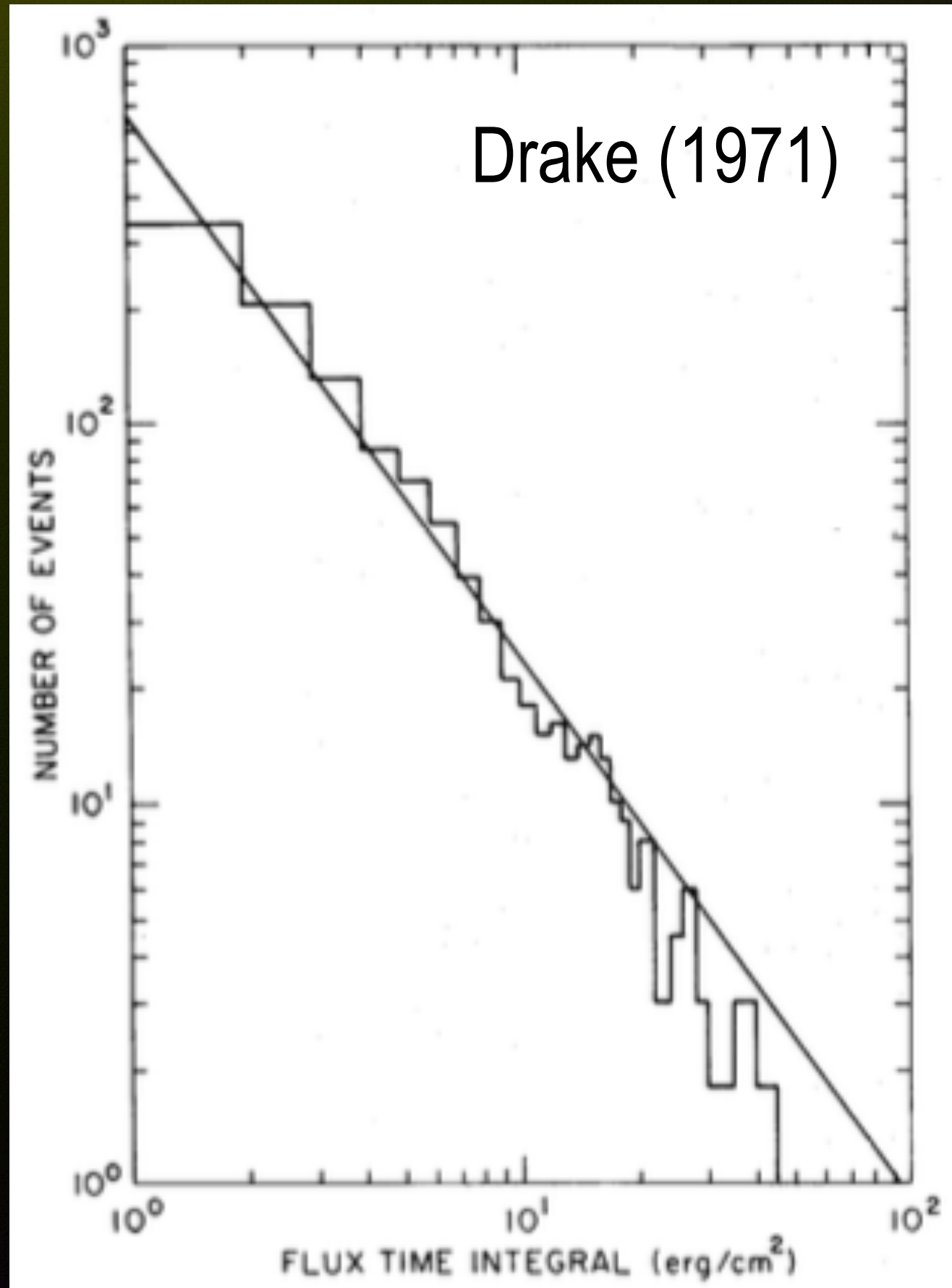
WHAT ARE SOLAR FLARES, PHYSICALLY AND STATISTICALLY?



Flare occurrence number
vs. integrated photon flux



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Flare occurrence number
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Flares are (Rosner & Vaiana 1978):

- Stochastic relaxation (storage and release) processes
- Physically uncoupled / independent
- Brief, comparing to intermediate times between flares

$$P(t) = \bar{\nu} e^{-\bar{\nu}t}$$

- Leading to a power-law occurrence frequency for flare energies

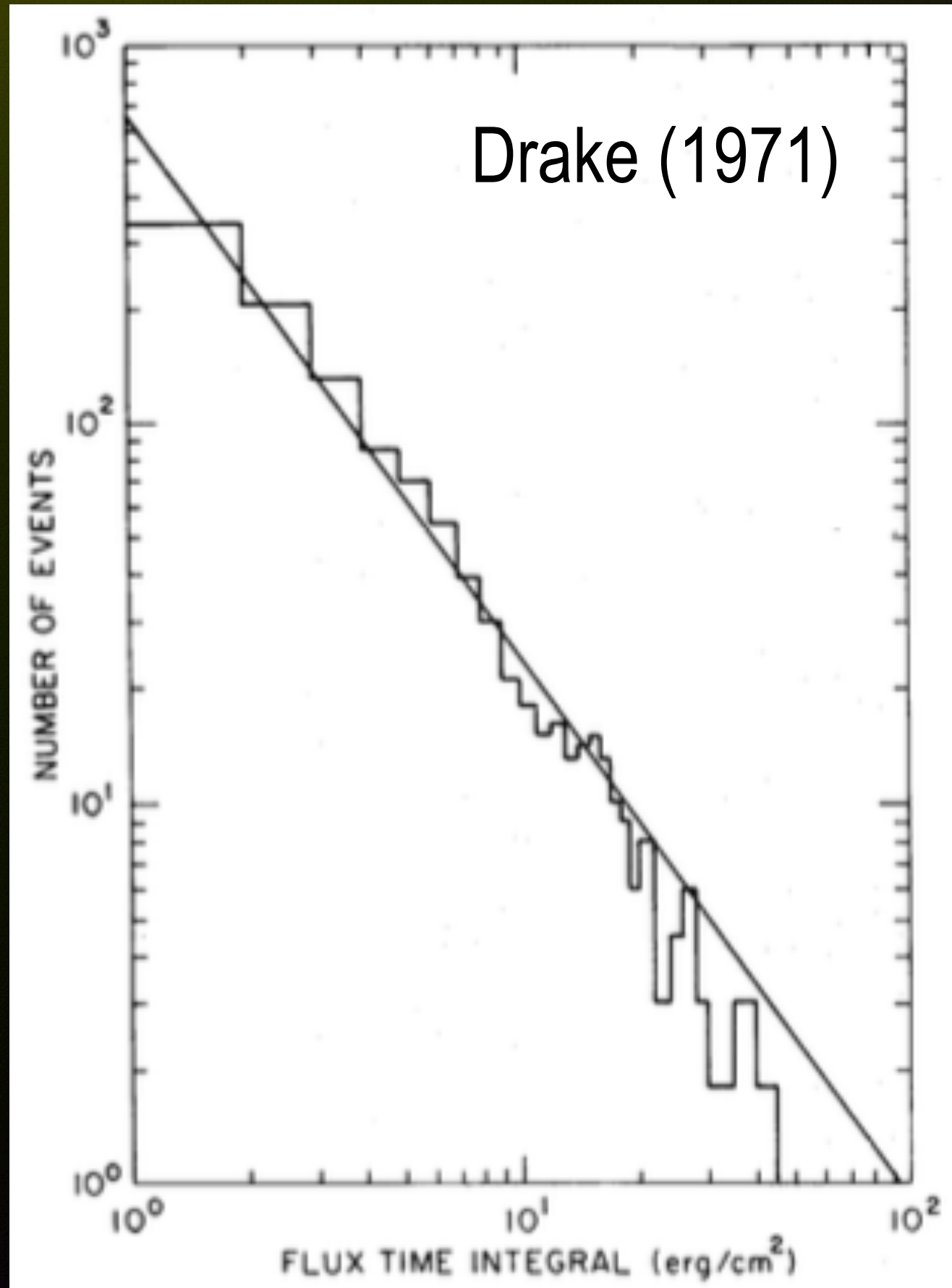
$$P(E) \sim \left(1 + \frac{E}{E_0}\right)^{-\gamma}$$

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Goa, January 29, 2016



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Flare occurrence number
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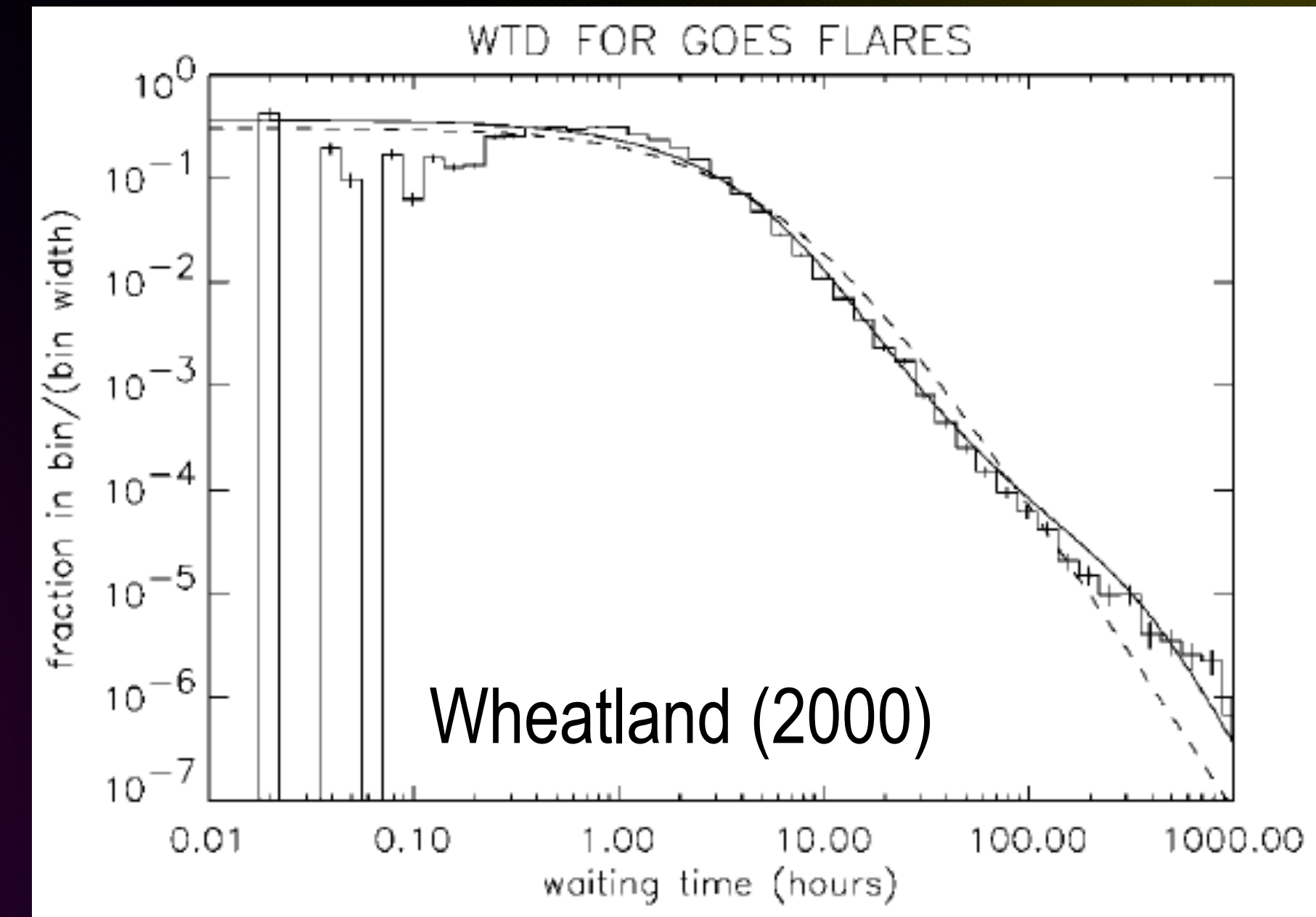
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Flare occurrence is a time-dependent Poisson process (also Wheatland & Litvinenko 2002). This can explain power laws in the flare distribution functions.

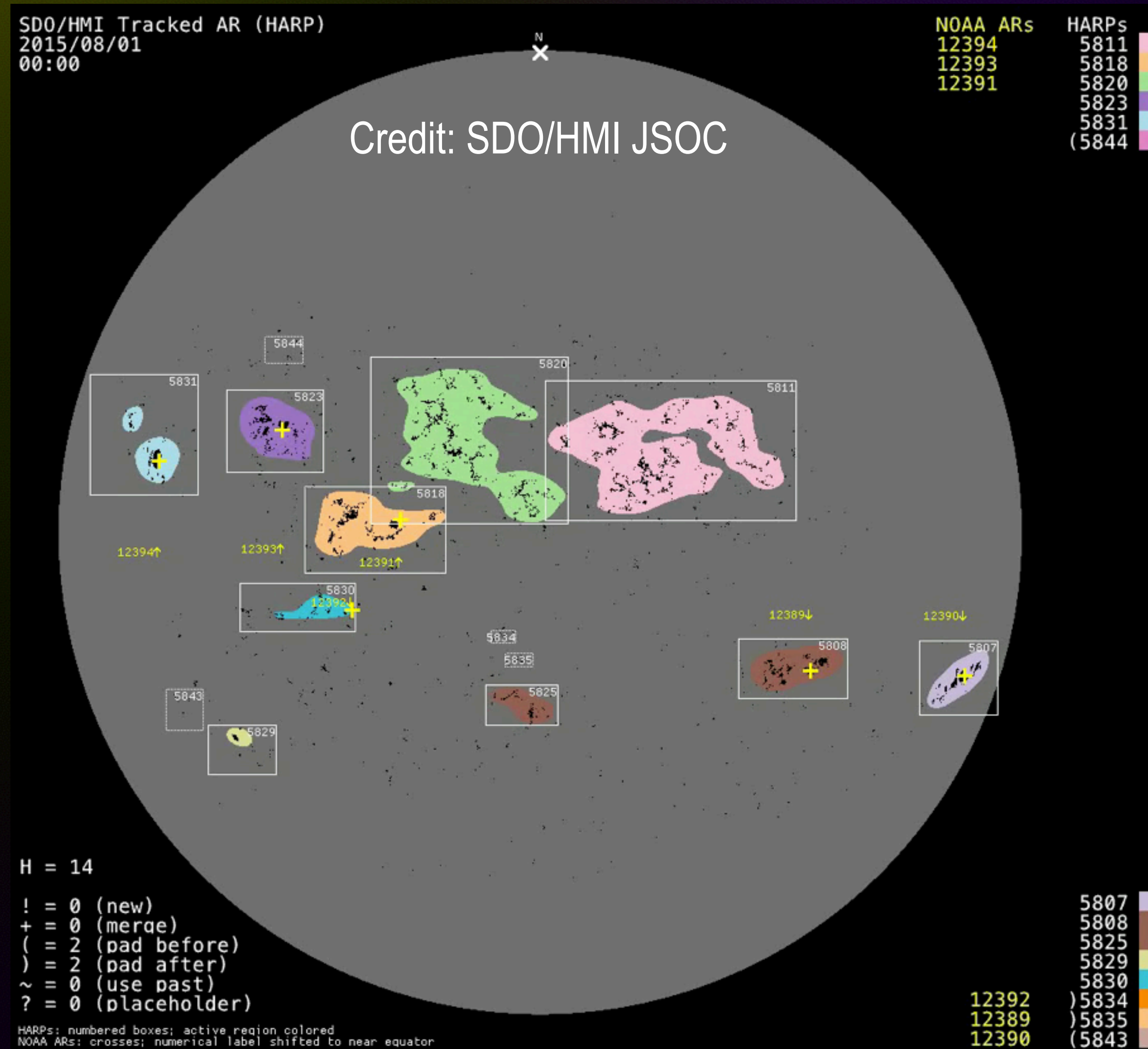
$$P(k) = \frac{\bar{\nu}^k}{k!} e^{-\bar{\nu}}$$

Gallagher et al. (2002)

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THE FLARE-PREDICTION CHALLENGE ...



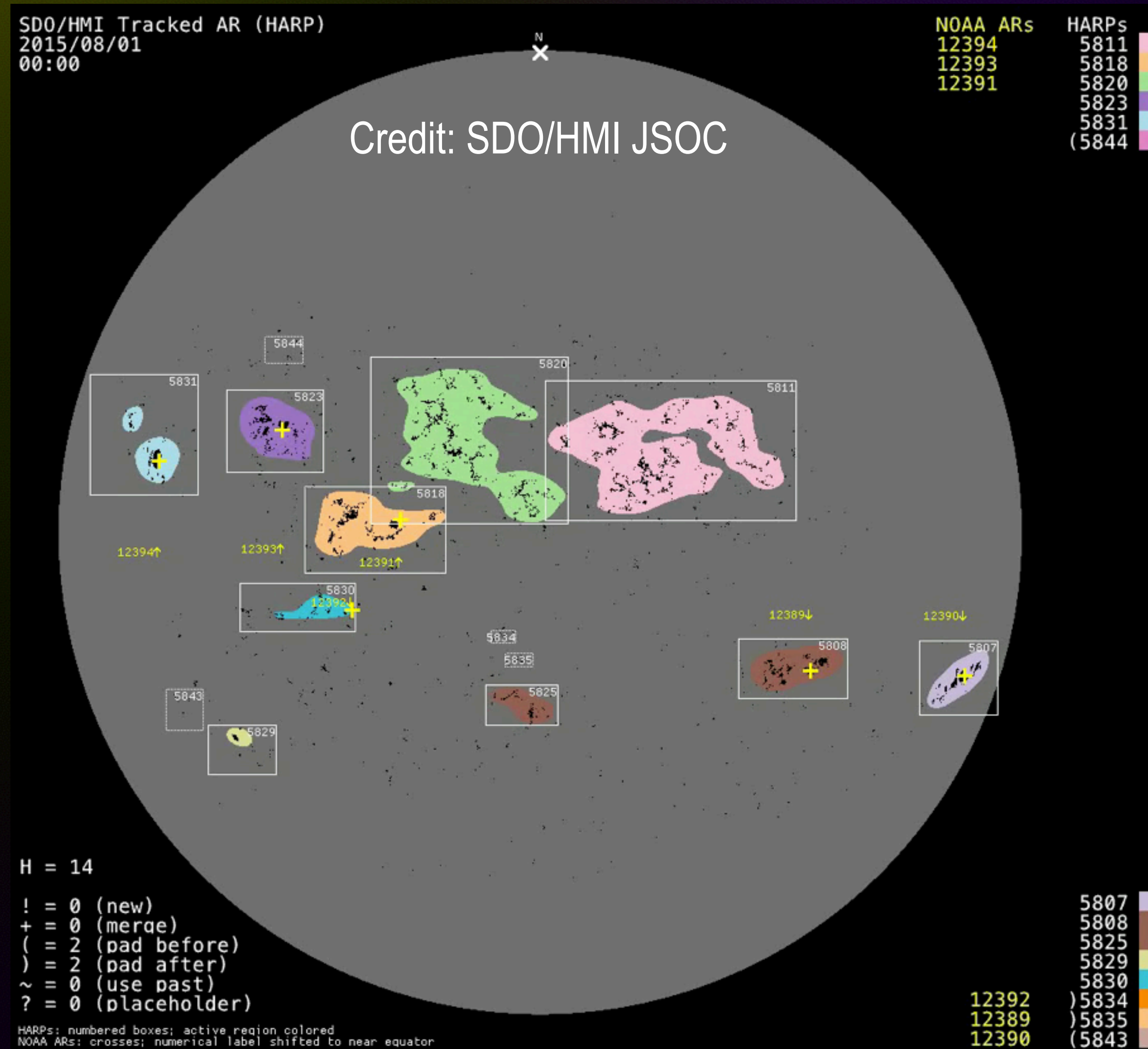
Objective: predict solar flares from near real-time (NRT) observations of solar evolution

- Prediction typically involves solar photospheric (LOS or vector) magnetic field measurements. SDO/HMI is the most prominent source of these data
- Predictive parameters are inferred locally
- Observational cadence:
 - 45 s, for full-disk LOS data
 - 720 s, for full-disk vector data

Reference: SOHO/MDI full-disk LOS data:
5,760 s



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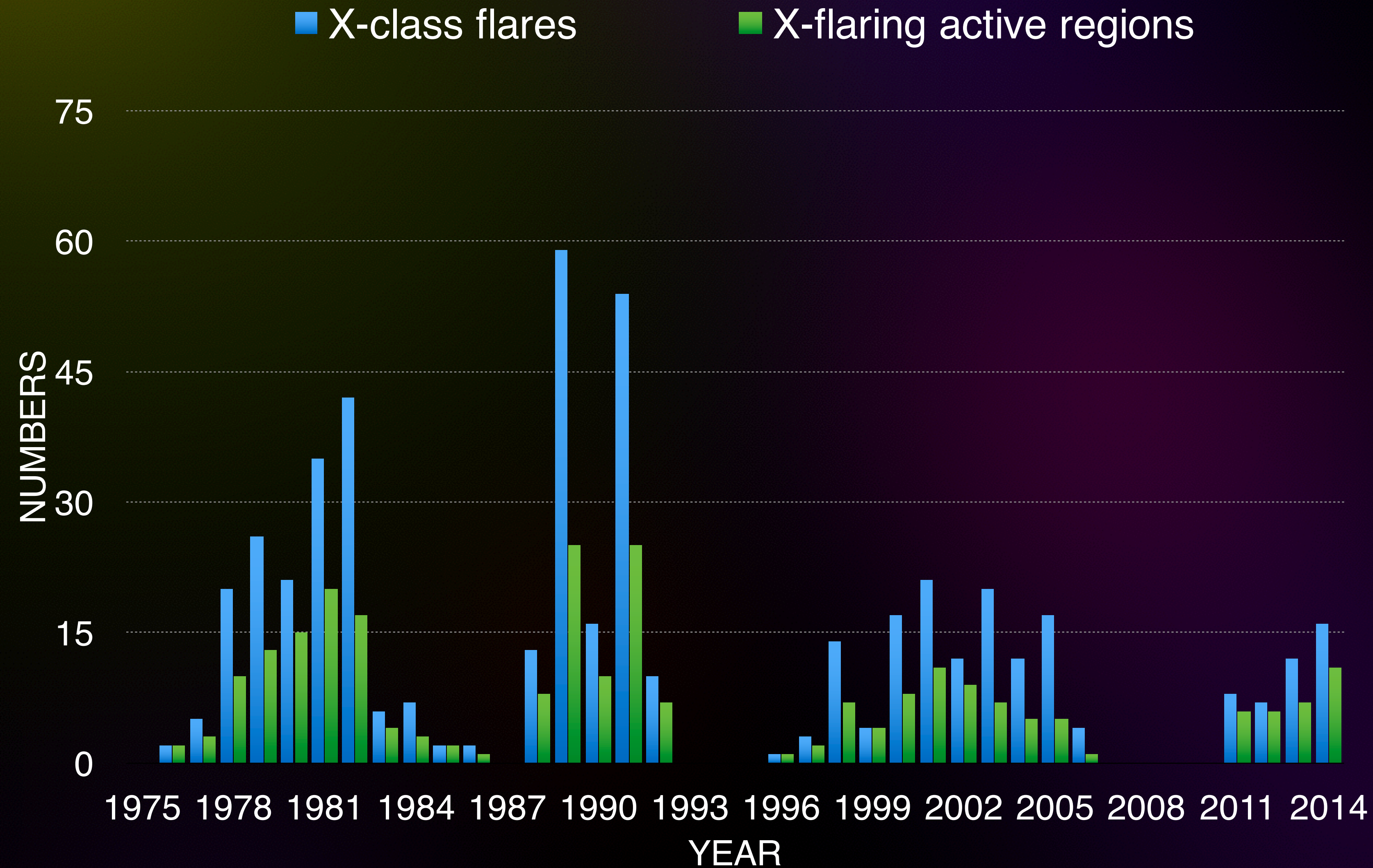
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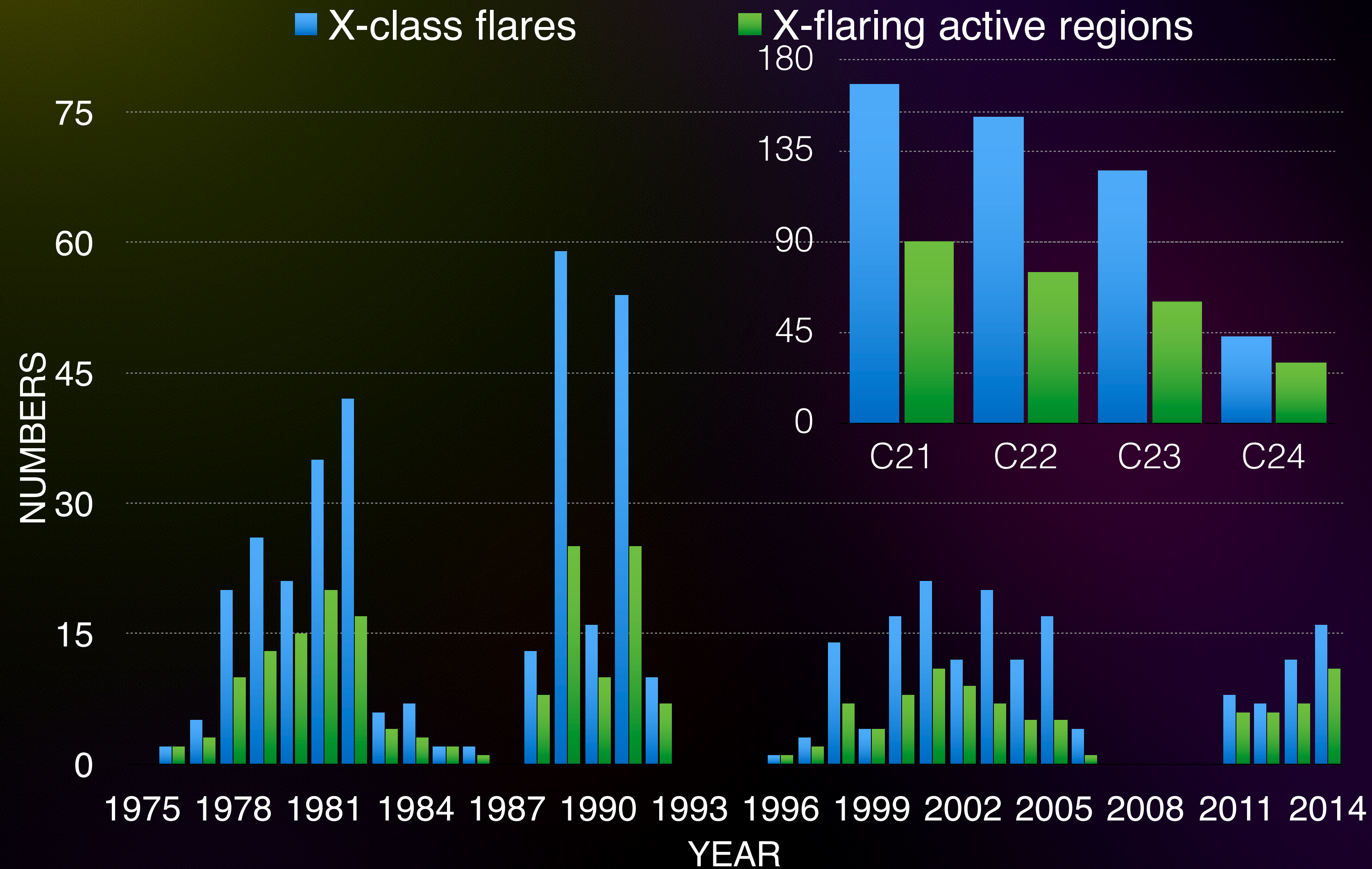
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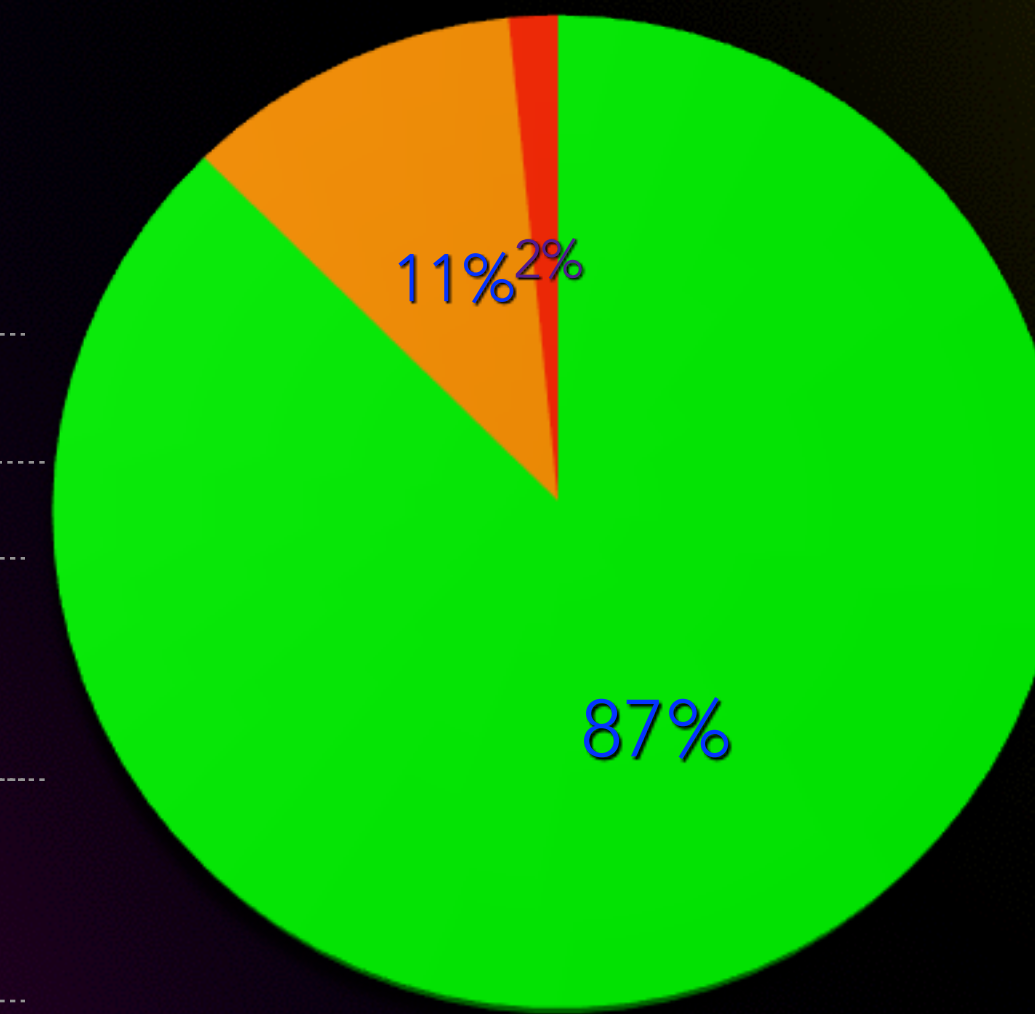
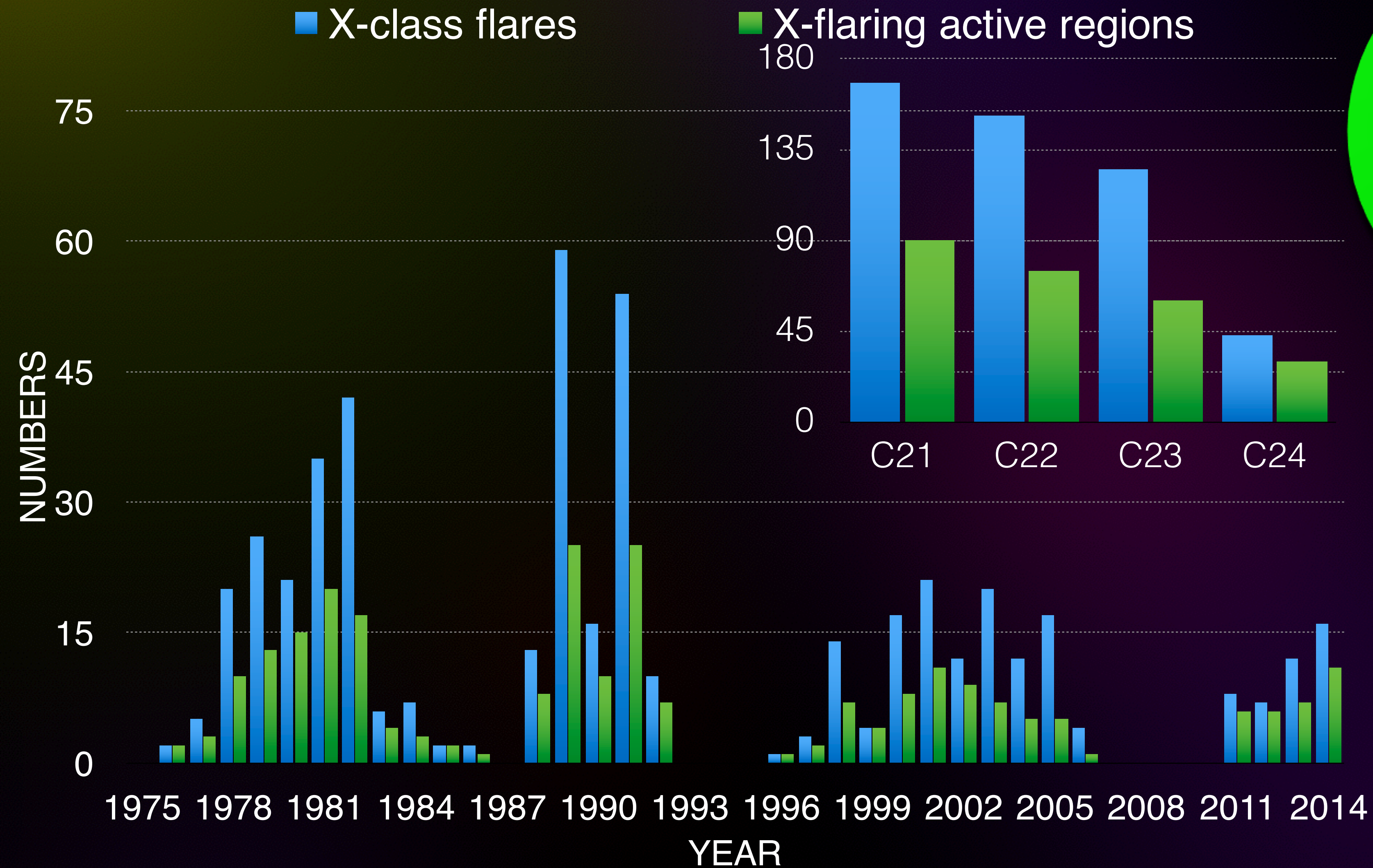
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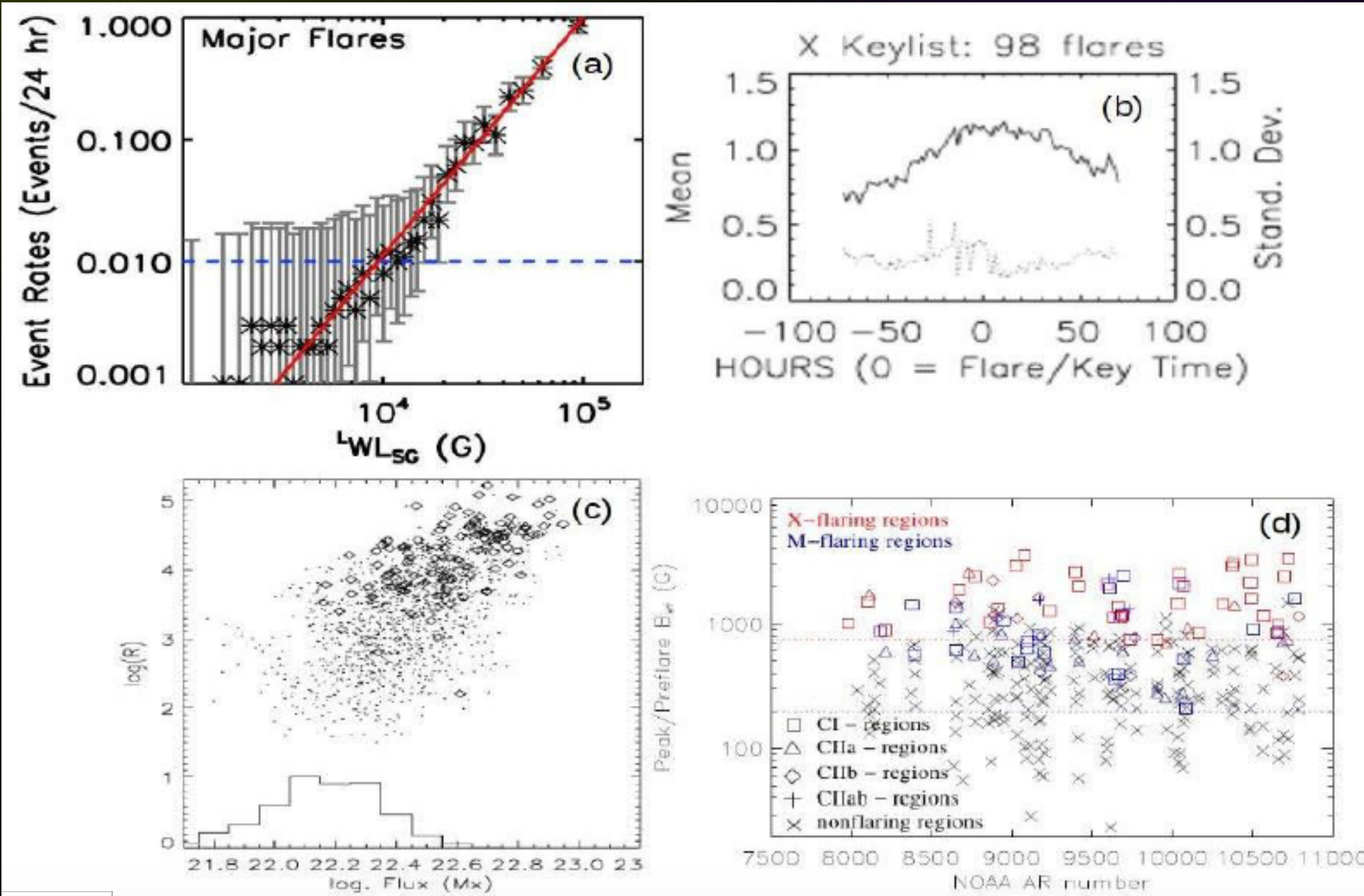
- Sub-flaring & C-class-flaring
- M-Class flaring
- X-class flaring

Only ~2% of active regions gave at least one X-class flare in C23!



DIFFERENT METHODS IN OPERATION WORLDWIDE

Mainly aiming to quantify the magnetic complexity of the host active regions



Upper left: W_{LG} - Falconer et al. (2011)

Upper right: GWILL 0- Mason & Hoeksema (2010)

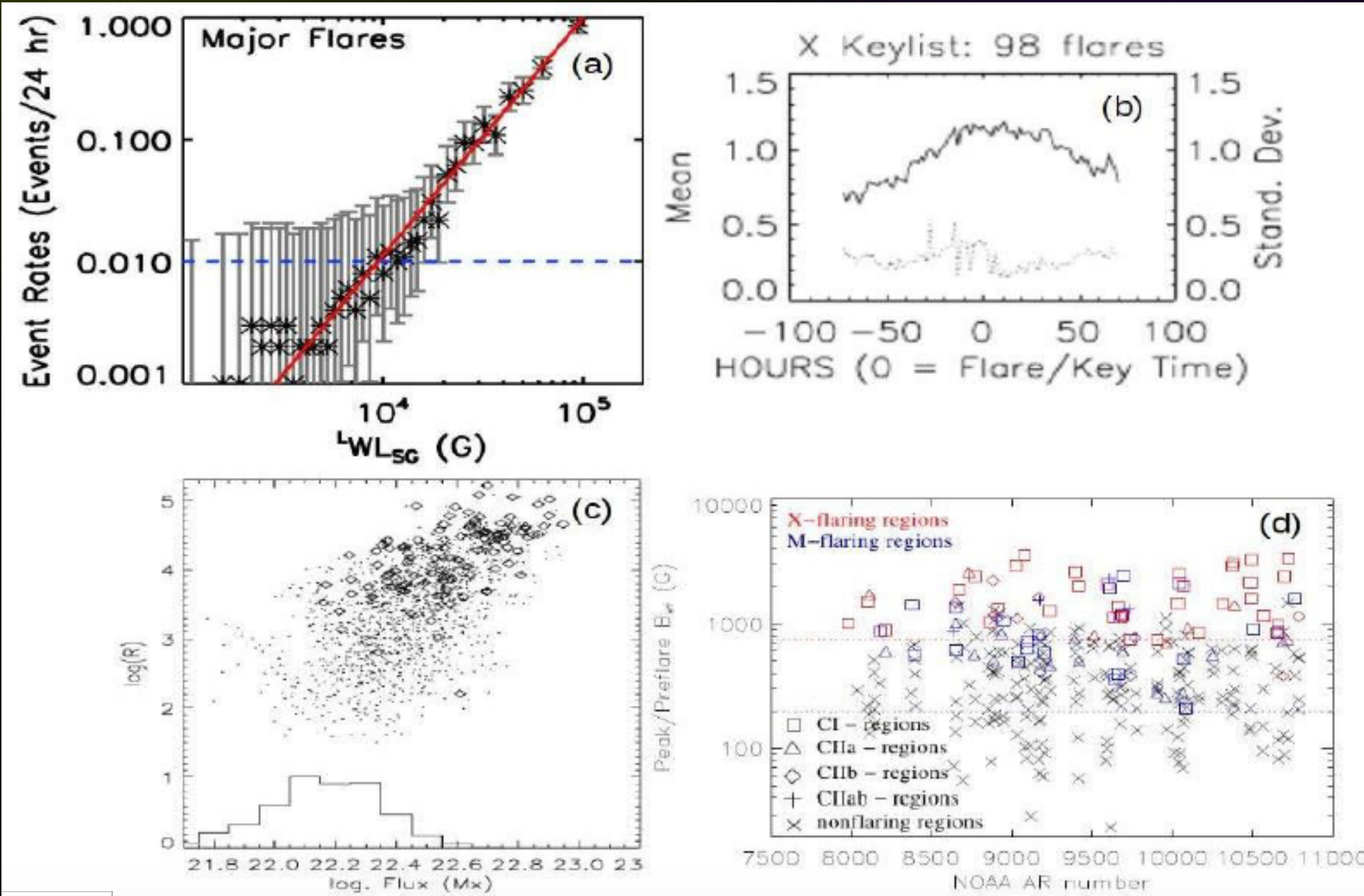
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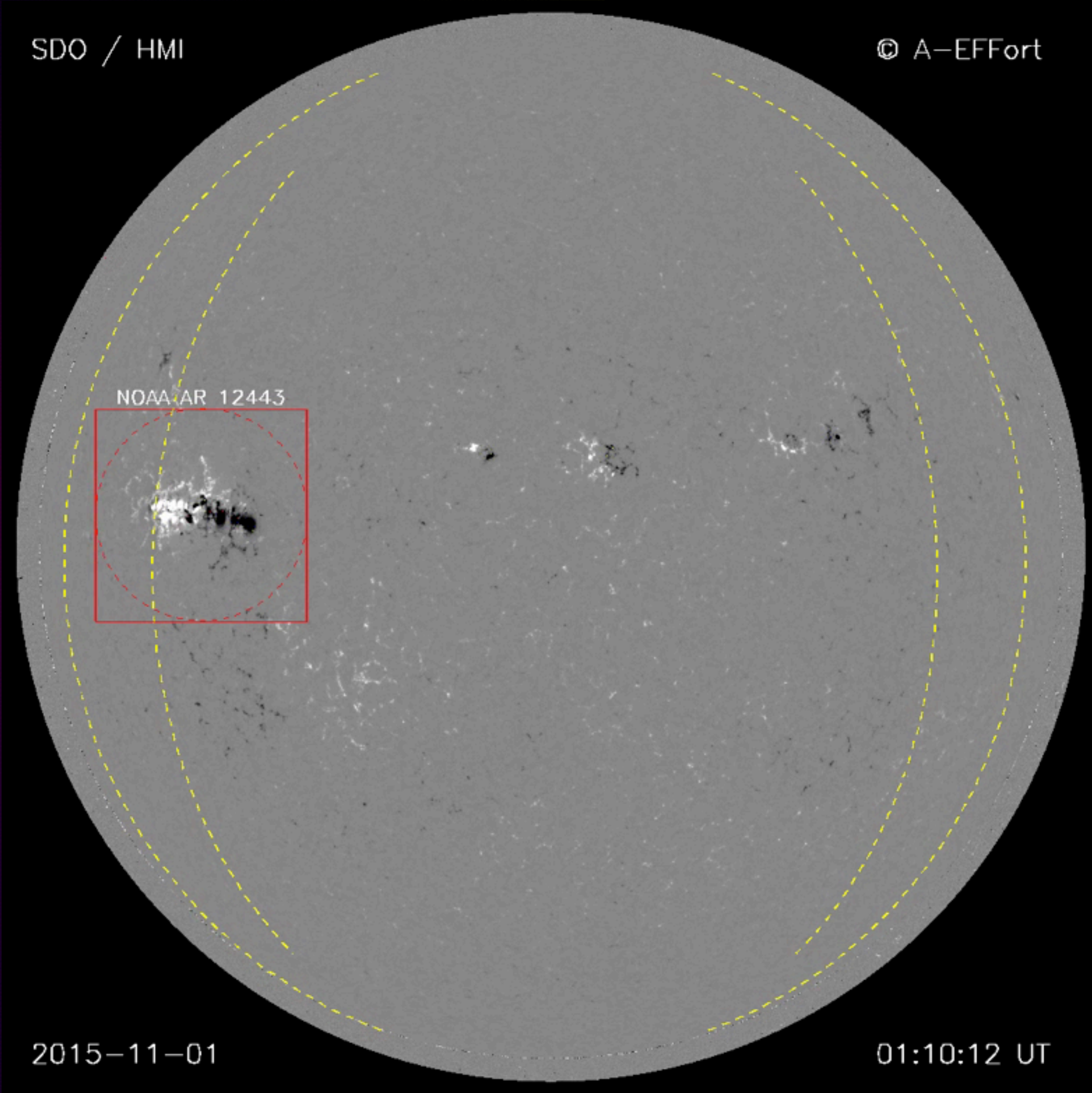
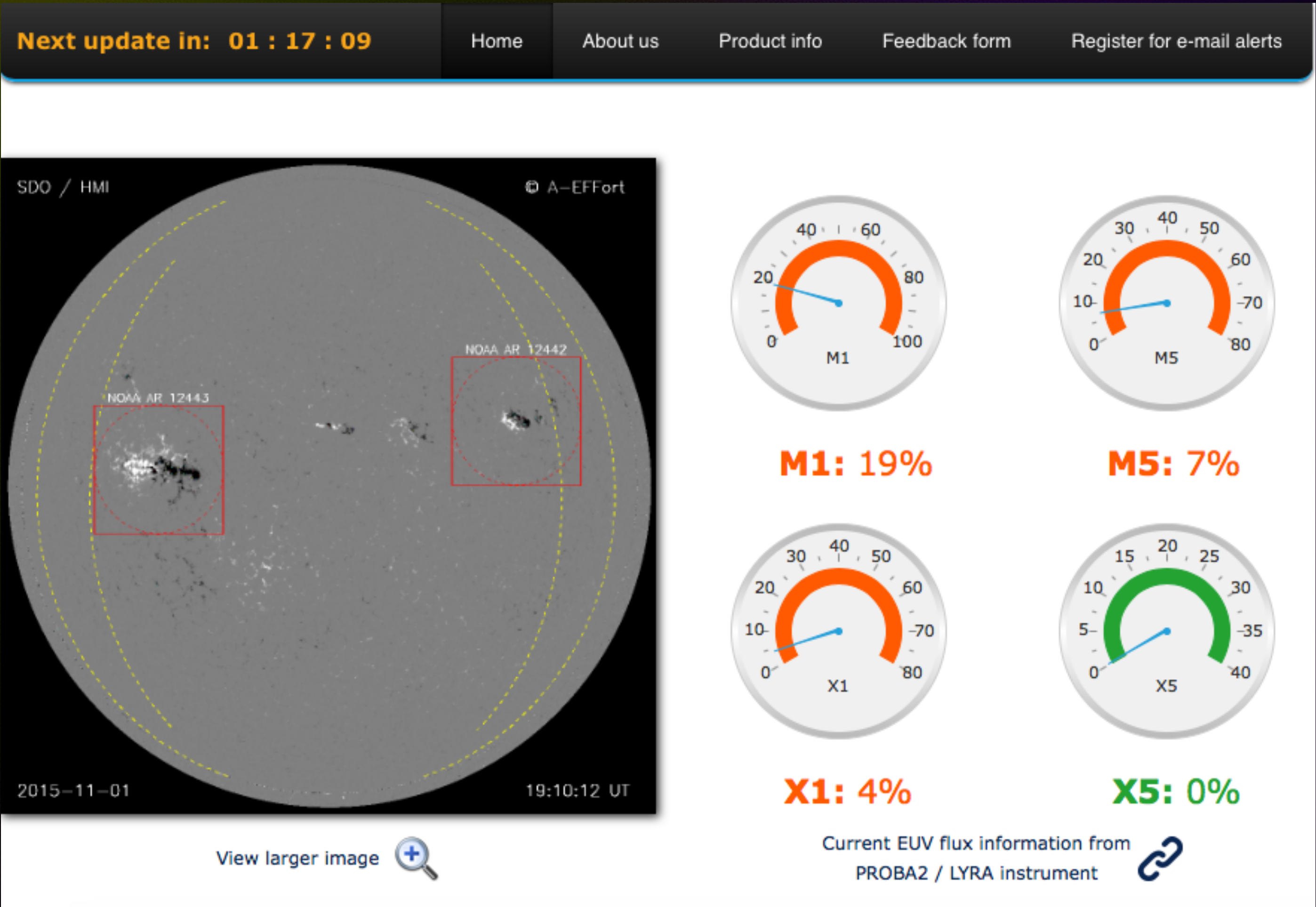
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Operational flare-prediction services:

- NOAA / SWPC
- Max Millennium / Solar Monitor
- UK Met Office
- U. Bradford / ASAP
- NASA / SRAG / MAG4
- ASSA / KSPC
- ESA / A-EFFort



ESA's SSA / SWE A-EFFort METHOD

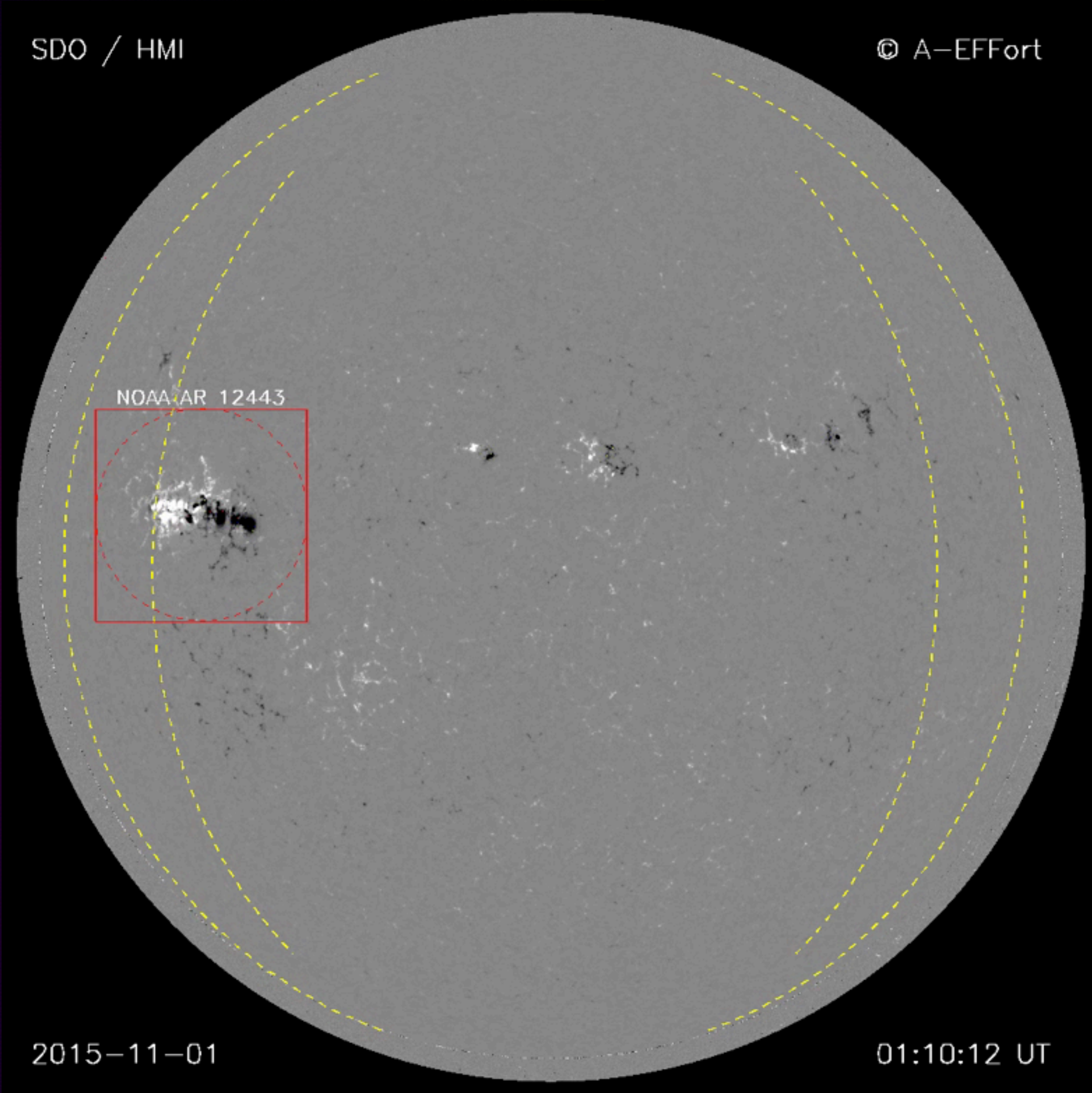
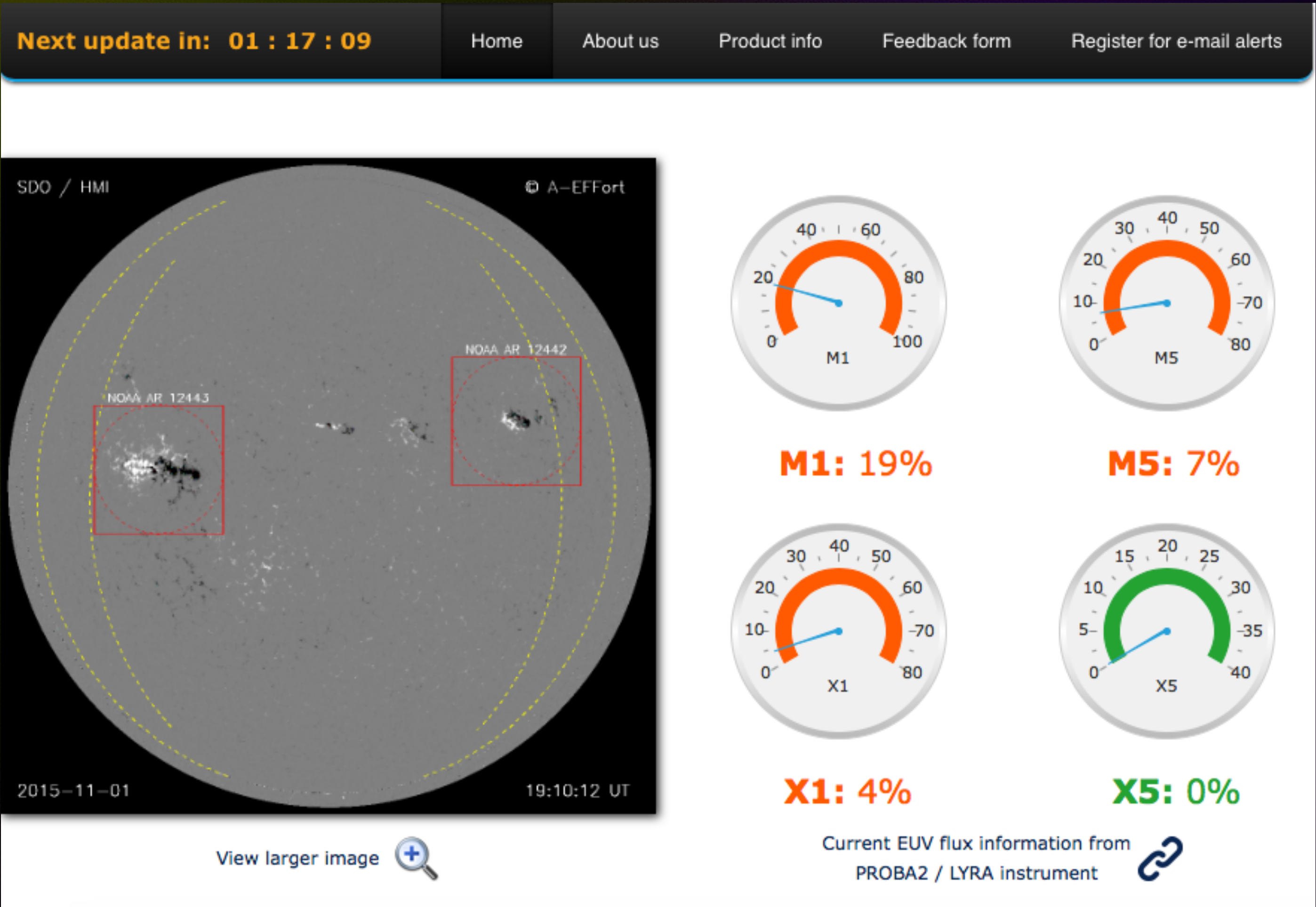


ESA's SSA/SWE A-EFFort interface

a-effort.academyofathens.gr



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IN FACT, MANY DIFFERENT METHODS...

Keyword	Description	Formula	F-Score	Selection
TOTUSJH	Total unsigned current helicity	$H_{c\text{total}} \propto \sum B_z \cdot J_z $	3560	Included
TOTBSQ	Total magnitude of Lorentz force	$F \propto \sum B^2$	3051	Included
TOTPOT	Total photospheric magnetic free energy density	$\rho_{\text{tot}} \propto \sum (B^{\text{Obs}} - B^{\text{Pot}})^2 dA$	2996	Included
TOTUSJZ	Total unsigned vertical current	$J_{z\text{total}} = \sum J_z dA$	2733	Included
ABSNJZH	Absolute value of the net current helicity	$H_{c\text{abs}} \propto \sum B_z \cdot J_z $	2618	Included
SAVNCPP	Sum of the modulus of the net current per polarity	$J_{z\text{sum}} \propto \left \sum_{B_z^+} J_z dA \right + \left \sum_{B_z^-} J_z dA \right $	2448	Included
USFLUX	Total unsigned flux	$\Phi = \sum B_z dA$	2437	Included
AREA_ACR	Area of strong field pixels in the active region	$\text{Area} = \sum \text{Pixels}$	2047	Included
TOTFZ	Sum of z-component of Lorentz force	$F_z \propto \sum (B_x^2 + B_y^2 - B_z^2) dA$	1371	Included
MEANPOT	Mean photospheric magnetic free energy	$\bar{\rho} \propto \frac{1}{N} \sum (B^{\text{Obs}} - B^{\text{Pot}})^2$	1064	Included
R_VALUE	Sum of flux near polarity inversion line	$\Phi = \sum B_{LoS} dA \text{ within } R \text{ mask}$	1057	Included
EPSZ	Sum of z-component of normalized Lorentz force	$\delta F_z \propto \frac{\sum (B_x^2 + B_y^2 - B_z^2)}{\sum B^2}$	864.1	Included
SHRGT45	Fraction of Area with shear > 45°	$\text{Area with shear} > 45^\circ / \text{total area}$	740.8	Included
MEANSHR	Mean shear angle	$\bar{\Gamma} = \frac{1}{N} \sum \arccos \left(\frac{B^{\text{Obs}} \cdot B^{\text{Pot}}}{ B^{\text{Obs}} B^{\text{Pot}} } \right)$	727.9	Discarded
MEANGAM	Mean angle of field from radial	$\bar{\gamma} = \frac{1}{N} \sum \arctan \left(\frac{B_h}{B_z} \right)$	573.3	Discarded
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MEANALP	Mean characteristic twist parameter, α	$\alpha_{\text{total}} \propto \frac{\sum J_z \cdot B_z}{\sum B_z^2}$	10.41	Discarded
TOTFX	Sum of x-component of Lorentz force	$F_x \propto - \sum B_x B_z dA$	6.147	Discarded
EPSY	Sum of y-component of normalized Lorentz force	$\delta F_y \propto \frac{- \sum B_y B_z}{\sum B^2}$	0.647	Discarded
EPSX	Sum of x-component of normalized Lorentz force	$\delta F_x \propto \frac{\sum B_x B_z}{\sum B^2}$	0.366	Discarded

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What is the optimal way to deal with all this information and still achieve reliable NRT forecasts within a preset forecast window?



VALIDATION : BORROWED BY TERRESTRIAL WEATHER FORECASTING

Categorical (dichotomous) validation: Flare (YES) or No Flare (NO)

	Forecast Flare	Forecast No-flare
Observed Flare	TP	FN
Observed No-flare	FP	TN

Table courtesy: Shaun Bloomfield

2 x 2 contingency table

- TP : true positives
- FN : false negatives
- FP : false positives
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- Generalized skill score:

$$SS = \frac{score - score_{reference}}{score_{perfect} - score_{reference}}$$



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- Heidke skill score (ref: random prediction):

$$HSS = \frac{2(TP + TN) - N}{N}$$

- Appleman skill score (ref: climatology [v]):

$$ApSS = \frac{TP - FP}{N}$$

- True skill statistic (ref: weighting POD w. POFD):

$$TSS = POD - POFD$$



CATEGORICAL FORECAST VERIFICATION METRICS

Metric Name	Short Name	Format	Worst Score	"No skill" Score	Perfect Score
Accuracy	ACC	$(TP + TN) / N$	0	...	1
Probability of detection	POD	$TP / (TP + FN)$	0	...	1
Probability of false detection (false alarm rate)	POFD	$FP / (FP + TN)$	1	...	0
False alarm ratio	FAR	$FP / (TP + FP)$	1	...	0
True skill statistic	TSS	$POD - POFD$	-1	0	1
Heidke skill score	HSS	$(TP + TN - E_{\text{random}}) / (N - E_{\text{random}})$	-1	0	1

Slide courtesy: Shaun Bloomfield



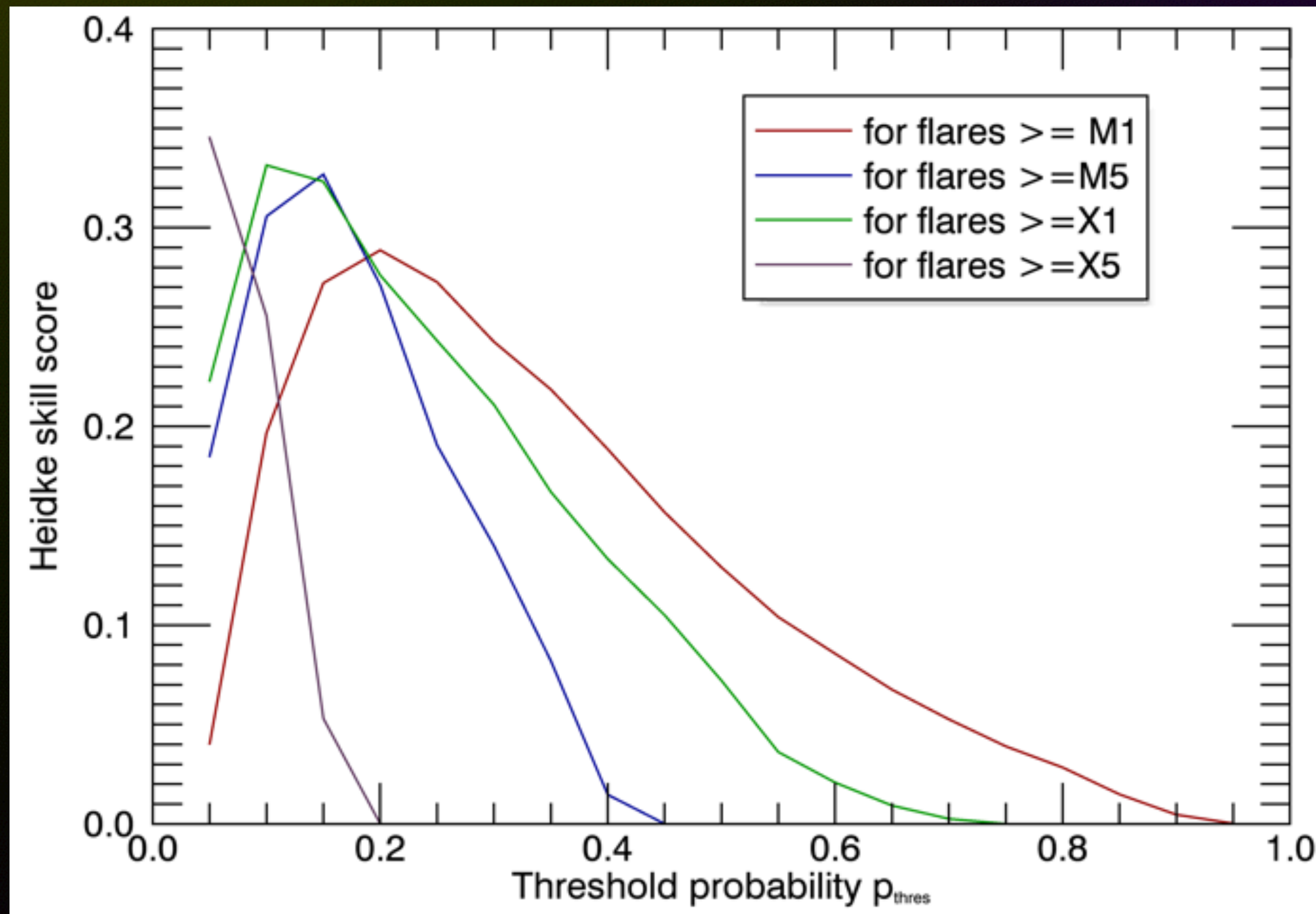
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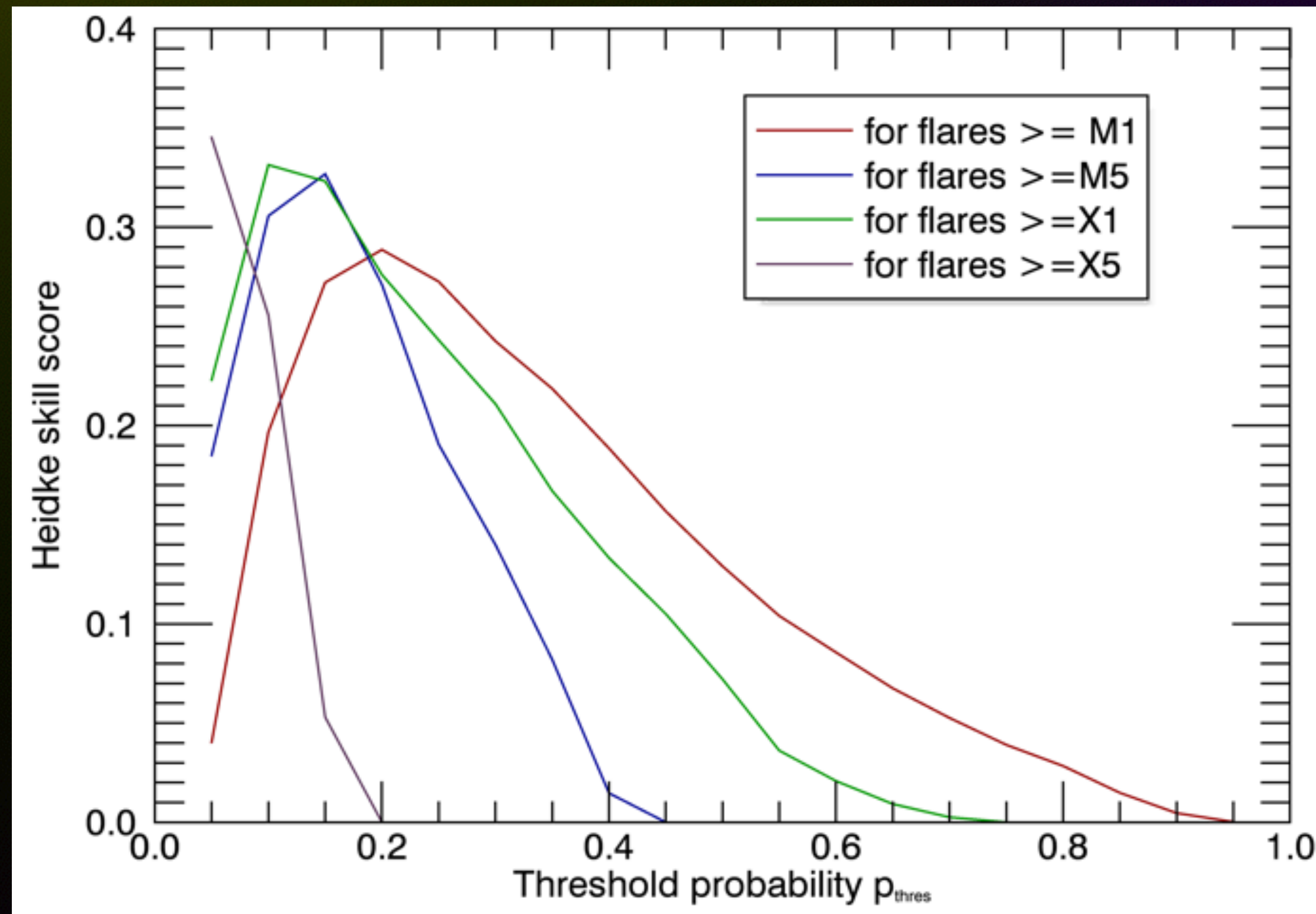


Example HSS vs. p_{thres} (A-EFFort)

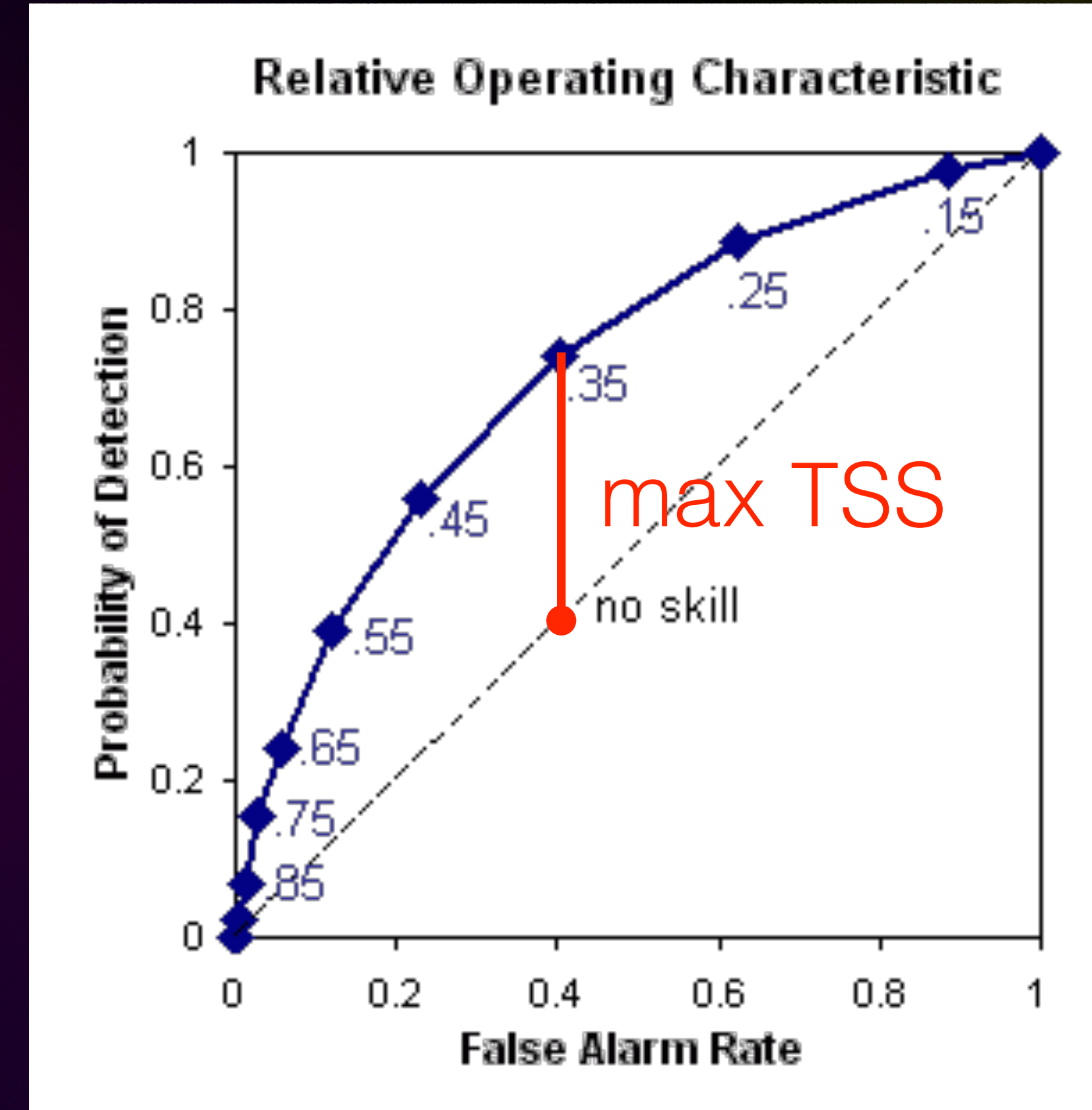


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Example HSS vs. p_{thres} (A-EFFort)



Example ROC curve (TSS vs. p_{thres})

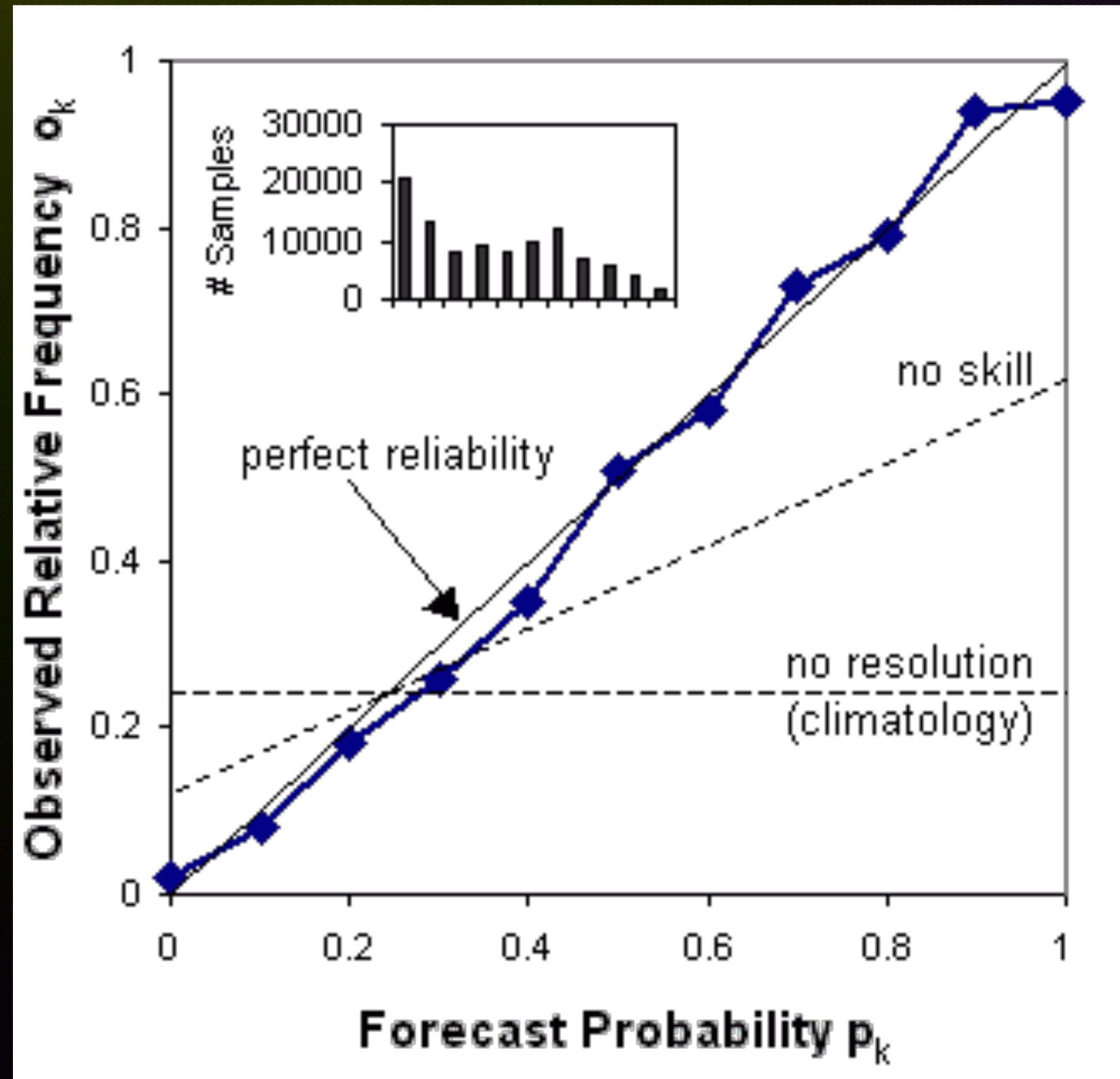
Source: WMO Forecast Verification Research

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PROBABILISTIC VALIDATION

Accept that a probability $0 < p < 1$ is assigned to each prediction



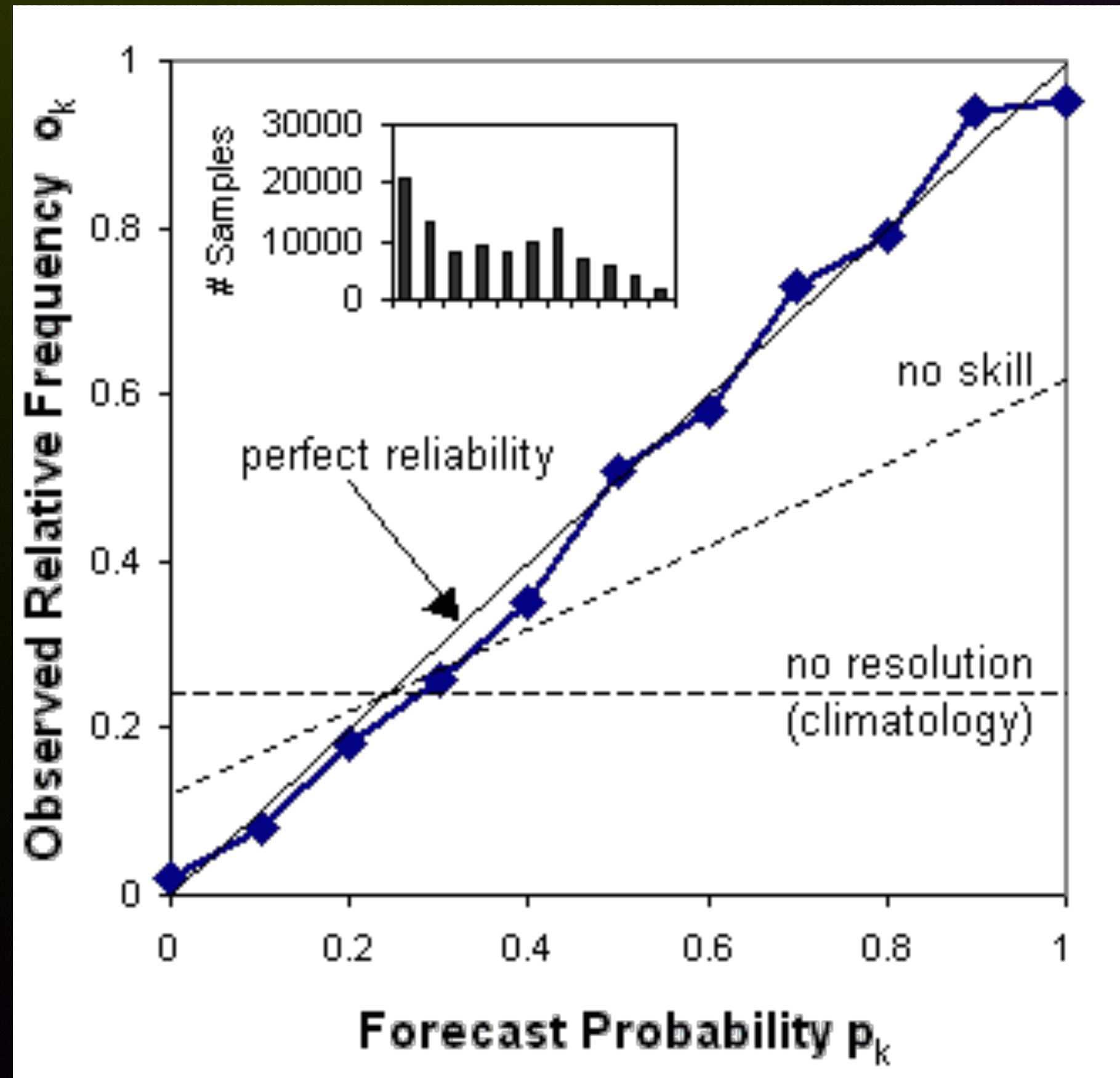
Reliability diagram

$$SS = 1 - \frac{MSE_{forecast}}{MSE_{reference}}$$



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Reliability diagram

- Correlate forecast probability with observed frequency
- Compare your skill against climatology (mean flaring rate within forecast window)
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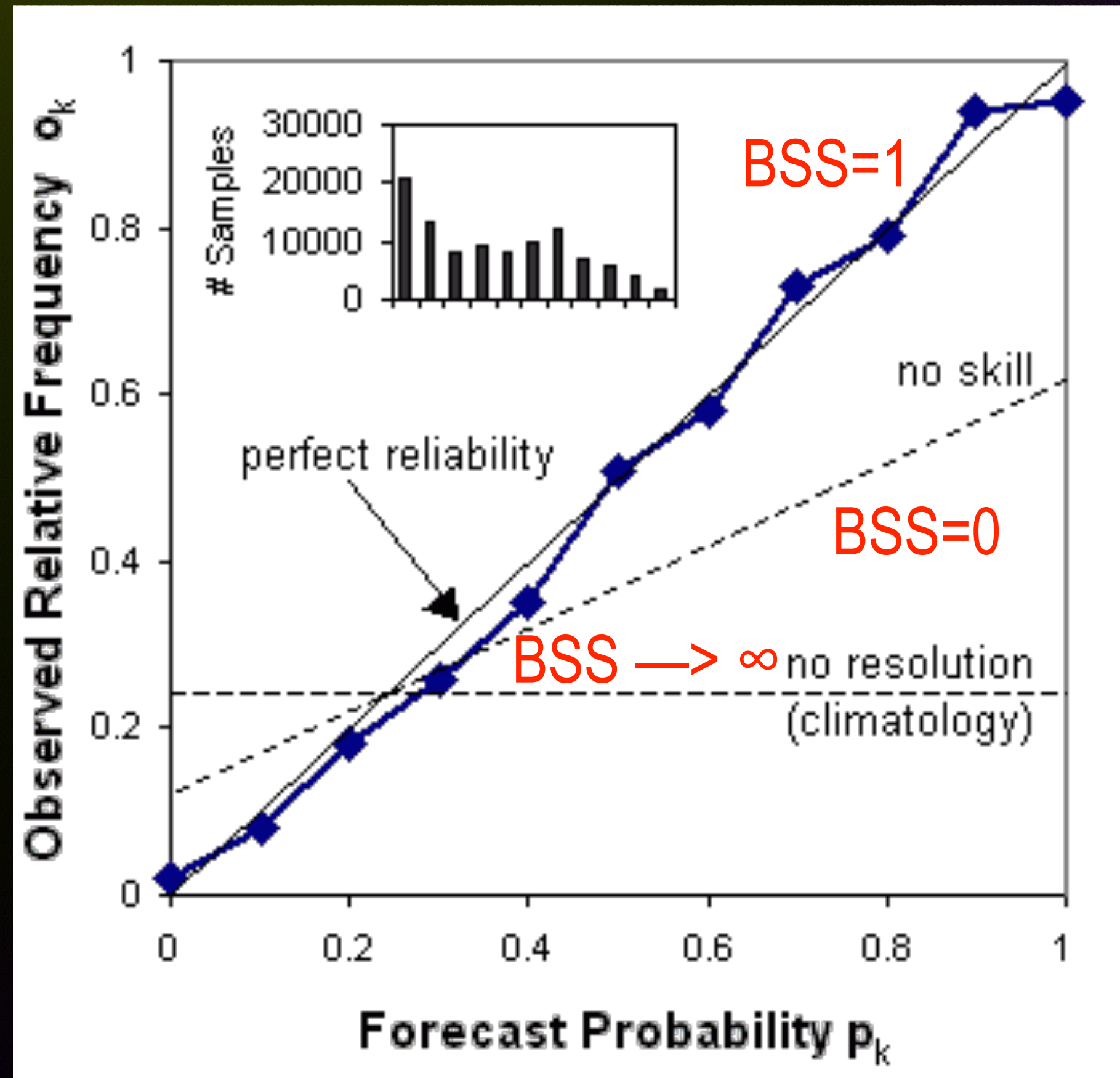
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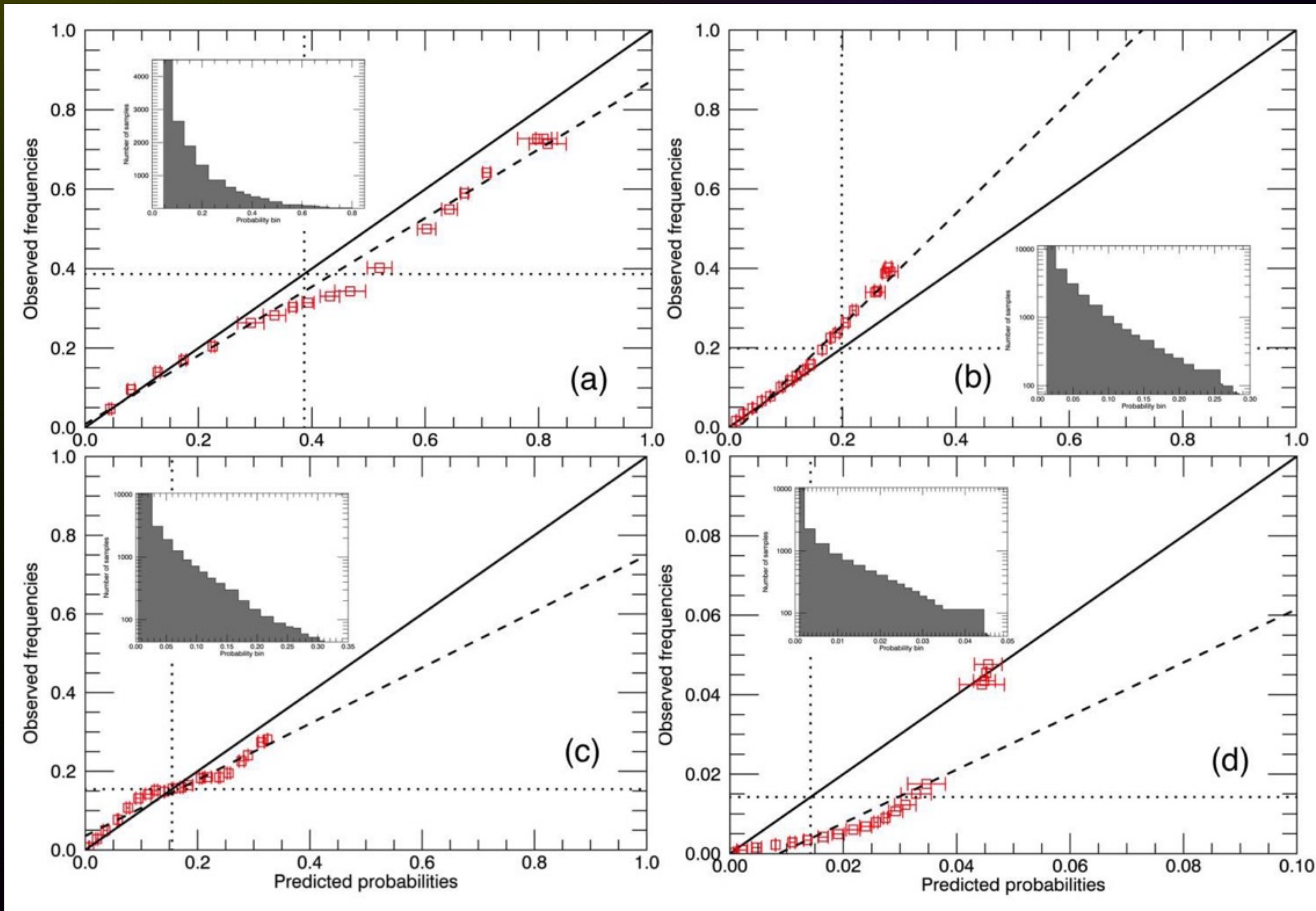
- Brier skill score (reference: climatology):

$$BSS = 1 - \frac{\langle (o - p)^2 \rangle}{\langle (o - \bar{o})^2 \rangle}$$

$$BSS \in (-\infty, 1)$$



PROBABILISTIC VALIDATION



Example probabilistic validation (A-EFFort):

(a) M1 and above: BSS = 0.88

(b) M5 and above: BSS = 0.78

(c) X1 and above: BSS = 0.80

(d) X5 and above: BSS = 0.38



TAILORING FLARE PREDICTION METHODS TO THE CUSTOMER NEEDS

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- Univariate forecasting is what all (but U. Bradford's ASAP) automated operational methods use
- Multivariate forecasting can also be used in the form of :
 - Multi-variable predictors:

$$\text{predictor} = \omega_1 \text{predictor}_1 + \omega_2 \text{predictor}_2 + \dots + \omega_n \text{predictor}_n$$

$$\omega_1, \omega_2, \dots, \omega_n \text{ unrestricted}$$

- Ensemble forecasting:

$$P(\text{flare}) = \omega_1 P_1(\text{flare}) + \omega_2 P_2(\text{flare}) + \dots + \omega_n P_n(\text{flare})$$

$$\sum_{i=1}^n \omega_i = 1$$



TAILORING FLARE PREDICTION METHODS TO THE CUSTOMER NEEDS

- Univariate forecasting is what all (but U. Bradford's ASAP) automated operational methods use
- Multivariate forecasting can also be used in the form of :
 - Multi-variable predictors:

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$$\omega_1, \omega_2, \dots, \omega_n \text{ unrestricted}$$

- Ensemble forecasting:

$$P(\text{flare}) = \omega_1 P_1(\text{flare}) + \omega_2 P_2(\text{flare}) + \dots + \omega_n P_n(\text{flare})$$

$$\sum_{i=1}^n \omega_i = 1$$

- **Task:** find $\omega_1, \omega_2, \dots, \omega_n$ such that validation results are optimized



TAILORING FLARE PREDICTION METHODS TO THE CUSTOMER NEEDS

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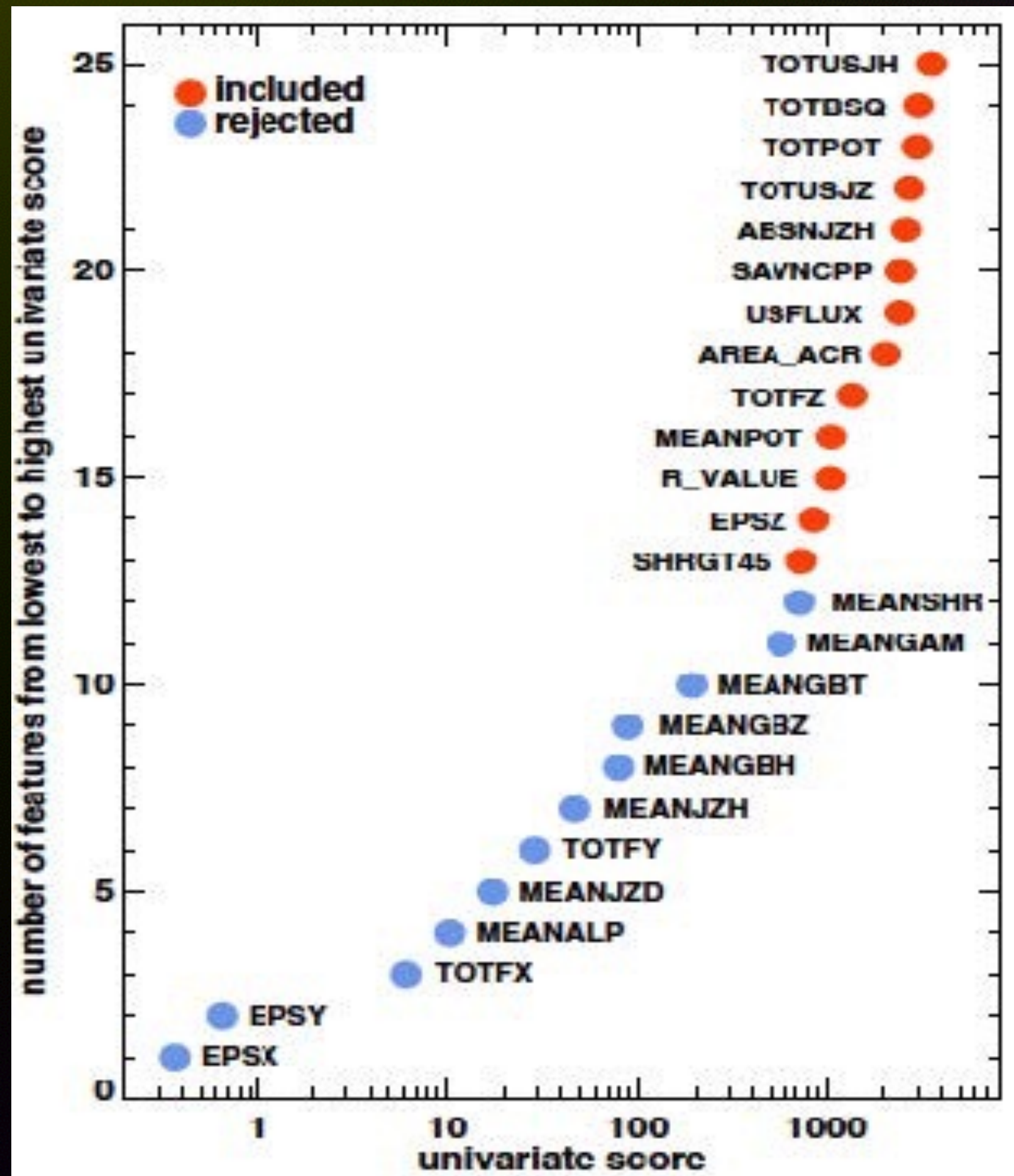
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However: optimization means different things to different customers!



SOME RECENT, PRELIMINARY EXAMPLES

- Multivariate forecasting



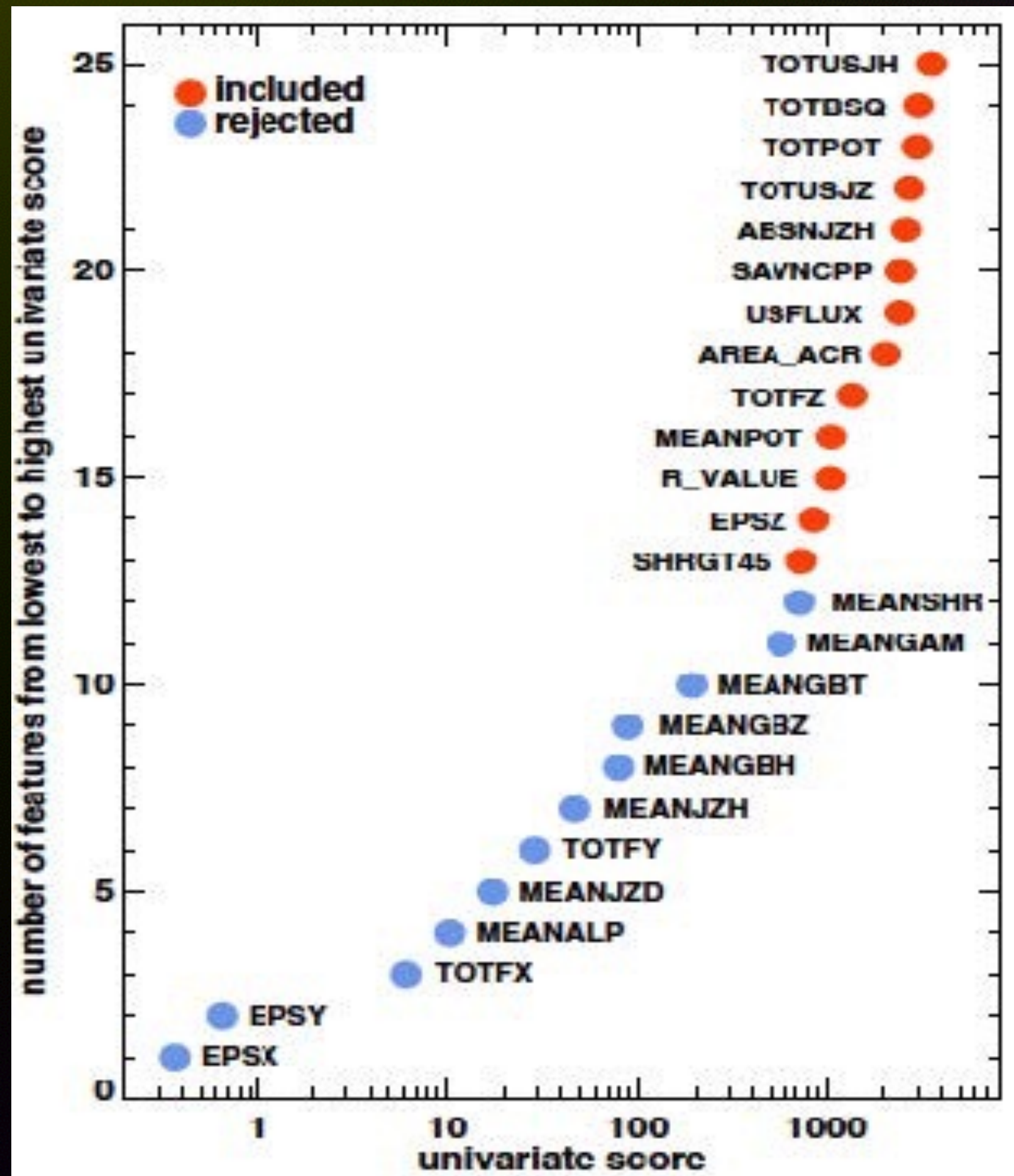
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Bobra & Couvidat (2015)



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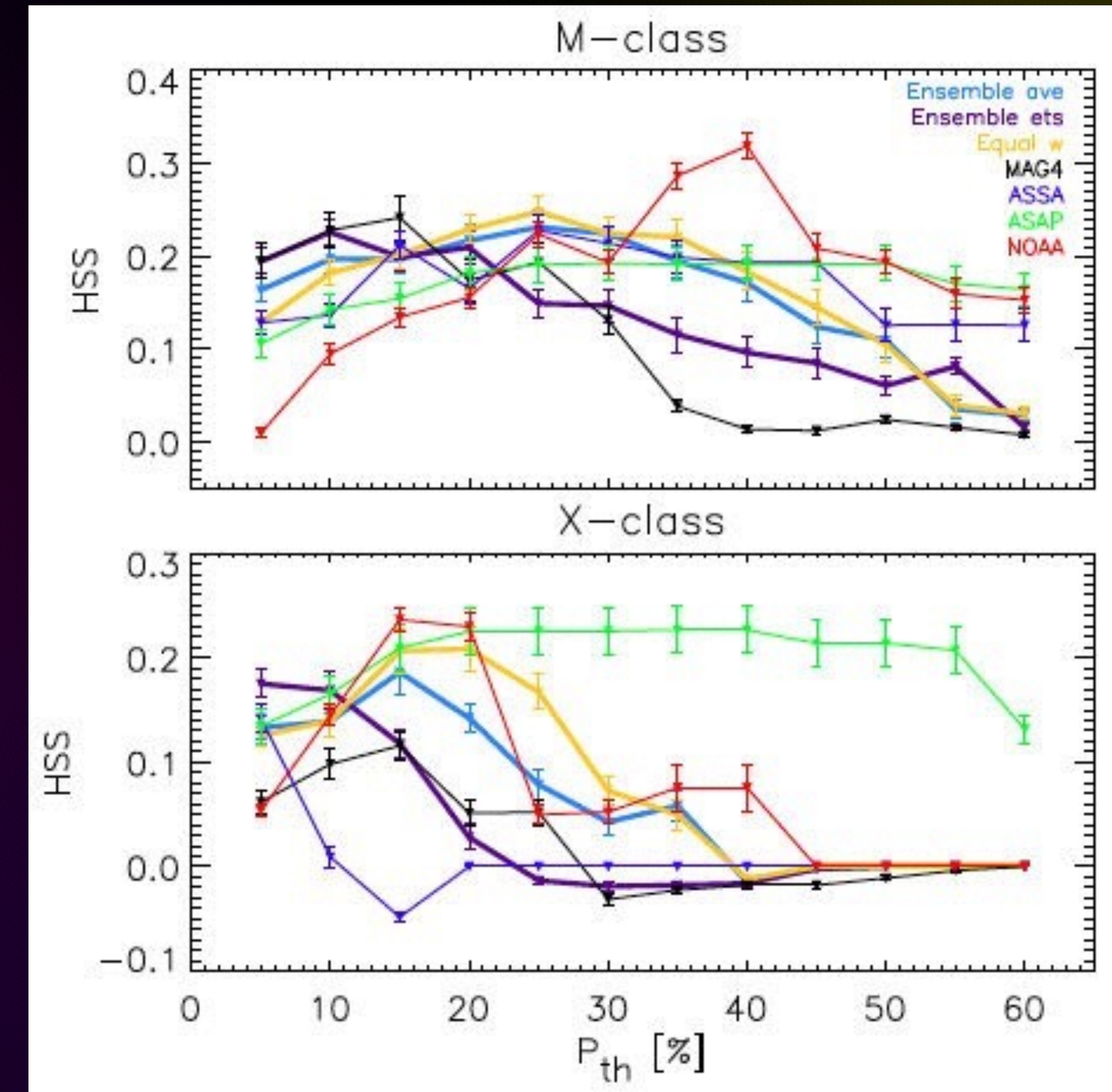
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Bobra & Couvidat (2015)

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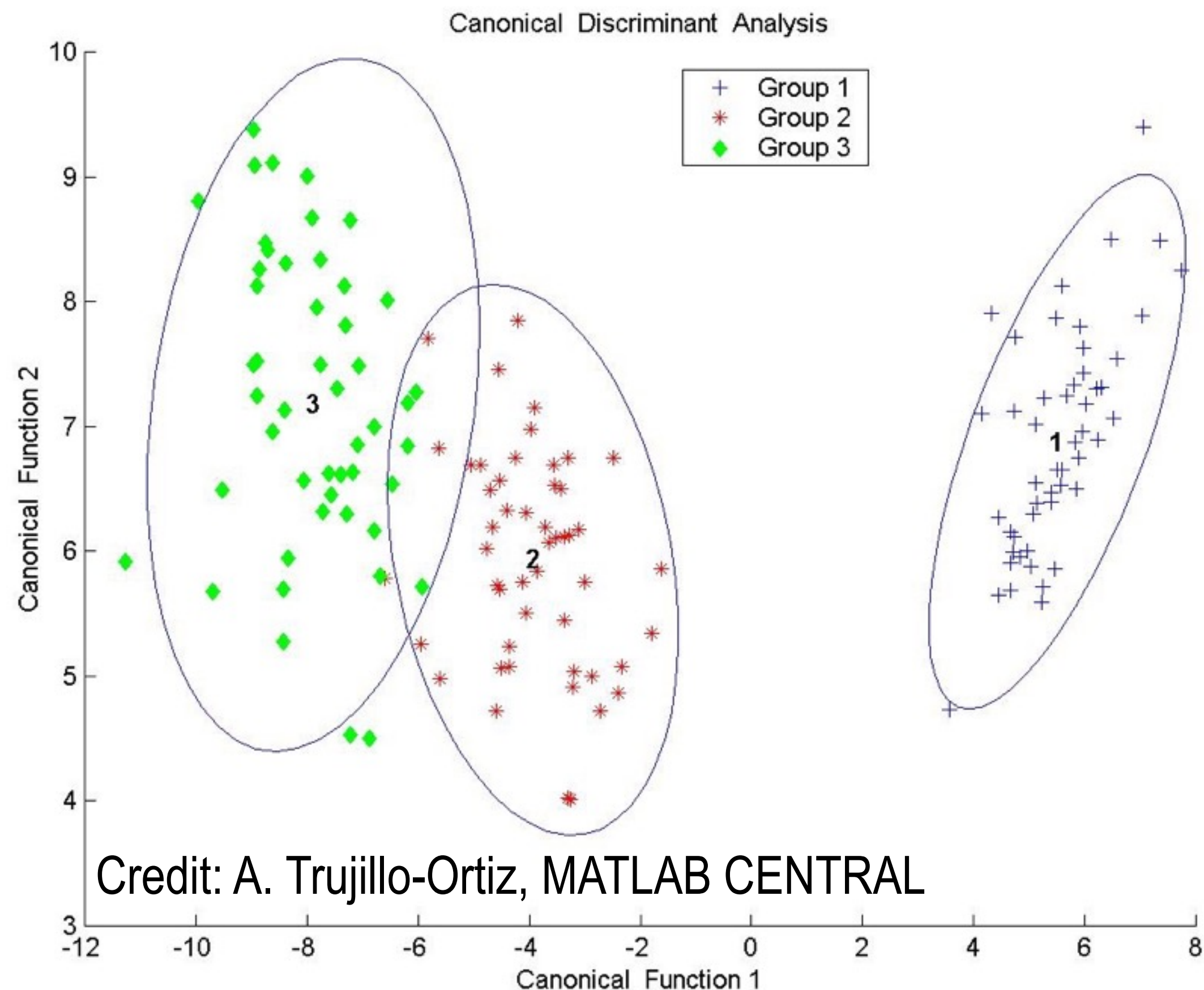
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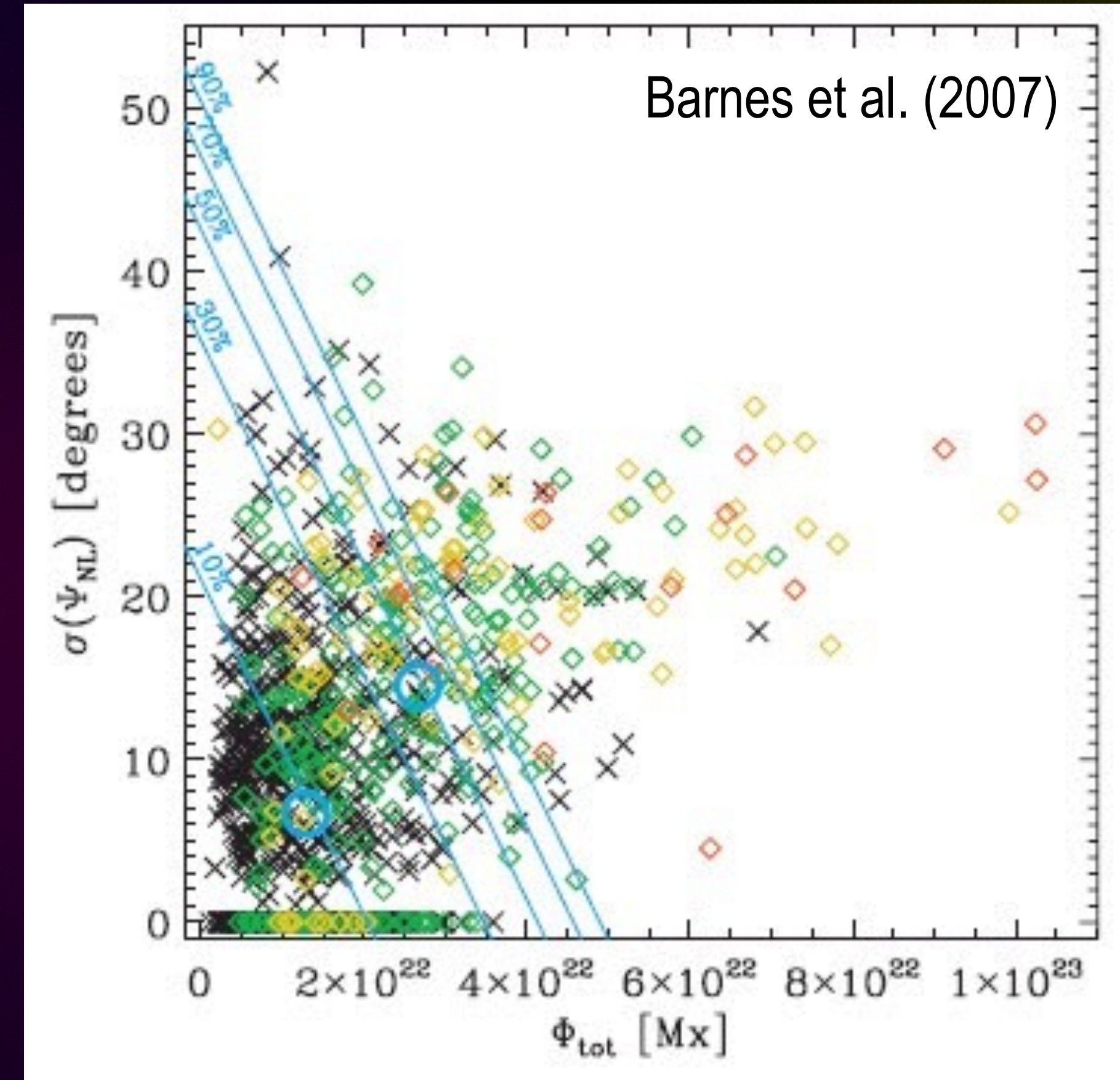
Guerra et al. (2015)



TAMING THE (LARGE) PARAMETER SPACE: DISCRIMINANT ANALYSIS and/or PCA



Example of two-function, three-group canonical DA



Two-function, linear DA for four-class prediction (non-flaring, C, M, and X-class)



FLARECAST: SYNTHESIS IN ACTION



FLARECAST is a European research project aiming to develop an automated solar-flare forecasting system with unmatched accuracy compared to existing facilities.

FLARECAST, using diverse European expertise, will:

- ★ use or reproduce already available predictors
- ★ classify predictors with respect to predictive ability
- ★ validate results in different ways

Project Partners:

- Academy of Athens, *Greece*
- Trinity College Dublin, *Ireland*
- Università Degli Studi Di Genova, *Italy*
- Consiglio Nazionale Delle Ricerche, *Italy*
- Centre National de la Recherche Scientifique, *France*
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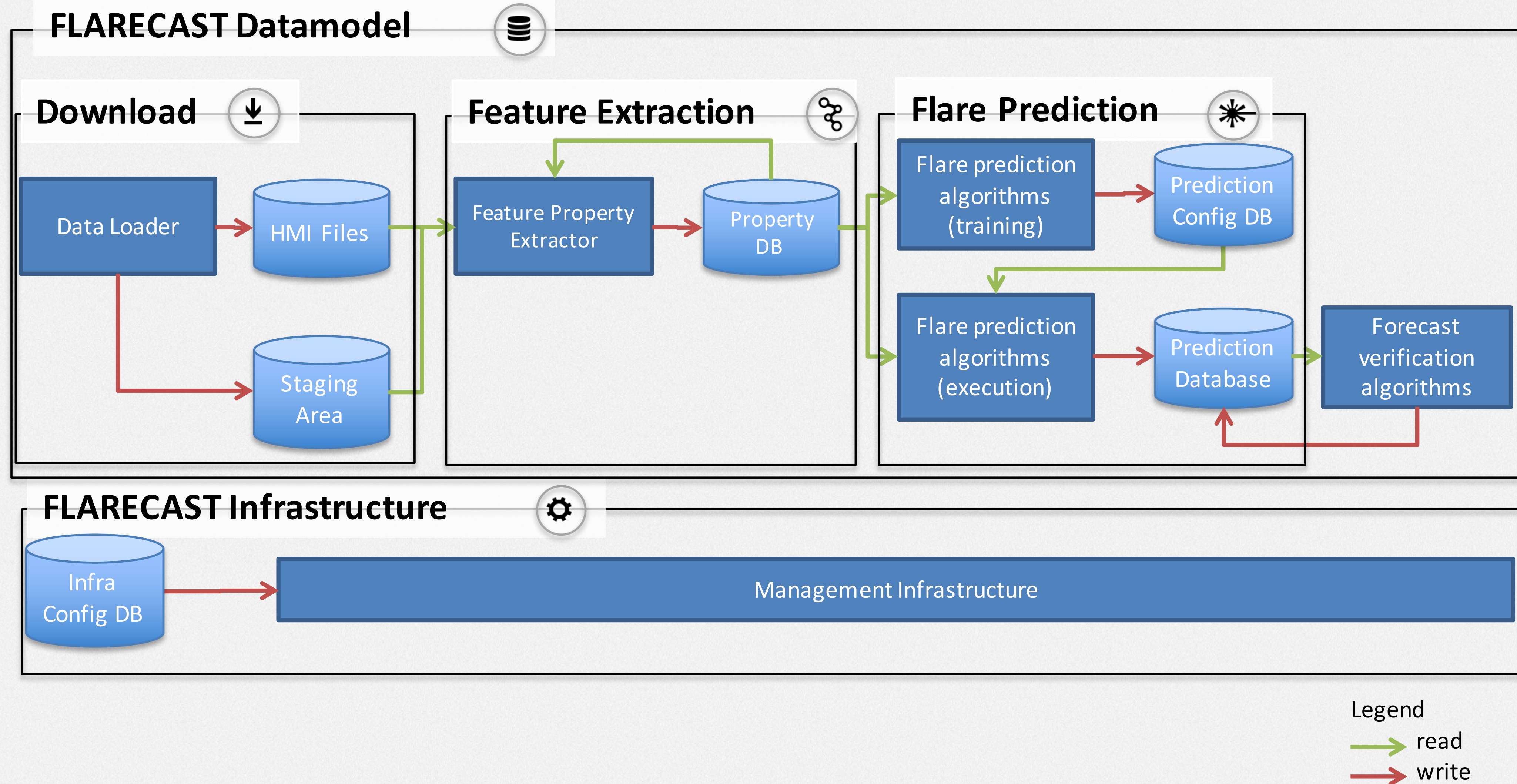
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The FLARECAST forecasting system will be openly accessible, featuring open-source software that will allow end users to perform their own tests. In this way FLARECAST will aim to both revamp solar flare prediction and contribute to a better understanding of the drivers of flare activity.



FLARECAST DATAMODEL & INFRASTRUCTURE : COLLECTIVE EXPERTISE



CONCLUSIONS

- ★ Solar flare prediction: consensus that it should be an asset of our space-weather forecasting toolbox
- ★ Validation: a vital task, done in multiple ways as per the customer's needs
- ★ We should understand how to enhance various validation metrics against others
- ★ Our ultimate task should be to bring a purely probabilistic prediction (due to stochasticity of the process) as close as possible to a categorical (YES / NO) one
- ★ The solution to this will not be unique - however, the used methods should be
- ★ Customized forecasts should rely on multivariate or ensemble modeling
- ★ Standard datasets could also be created for the validation of all methods



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All these tasks are being investigated by the FLARECAST Consortium



BACKUP SLIDES