



European Commission Horizon 2020 European Union funding for Research & Innovation

Solar Magnetic Data Analysis for the FLARECAST project

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Irish National Astronomy Meeting (INAM), UCD, Sept. 7 – 9 2016

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The FLARECAST Project

 Flare Likelihood And Region Eruption foreCASTing: EC H2020 consortium project.



Goals:

- Understanding properties and evolution that leads to flares.
- Enhance quality of flare predictions by using a great number of predictors and state-ofthe-art prediction algorithms.
- Allow end-users access to past and future data for scientific and operational purposes.

• Rapid enhancement of emission.



- Rapid enhancement of emission.
- EM nature: energy released via magnetic reconnection of B lines.

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Coronal loops

Reconnection of field lines

- Rapid enhancement of emission.
- EM nature: energy released via magnetic reconnection of **B** lines.
- Origin in strong-**B** areas in the Sun's surface (active regions)

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- Rapid enhancement of emission.
- EM nature: energy released via magnetic reconnection of **B** lines.
- Origin in strong-**B** areas in the Sun's surface (active regions)
- Classified according to peak in X-rays near Earth.



Photospheric predictors of Solar Flares

• Search for patterns in the Sun's photosphere and empirical relations.

Magnetic Field or Magnetogram

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Visible surface



Magnetic class (Bornmann & Shaw, 1994) Length of Polarity Inversion Line (PIL, Falconer 2003) Total magnetic flux McIntosh class (1990) Sunspot group area (Qahwaji & Colak, 2008)

Work Package 2: Active regions properties

- TCD focuses on the satellite data analysis for extracting AR properties (WP2).
- These properties will be used by the prediction algorithms.
- Calculate as many as possible (relevant) parameters.

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Property Extraction Algorithms

LOS magnetograms	
	SMART-derived properties (Ahmed et al., 2013)
	SMART delta finder (Padinhatteeri et al., 2015)
	Effective connected magnetic field strength (B _{eff}) (Georgoulis & Rus 2007)
	Fractal dimension (Georgoulis, 2012)
	Multi-fractal structure function s(q) inertial range index k (Georgouils, 2012)
	Fourier power spectral index (Guerra et. al., 2015)
	CWT power spectral index (Hewett et. al., 2008)
	Generalised correlation dimension (Georgoulis, 2012)
	Holder exponent h (Conlon et al., 2010)
	Hausdorff dimension D(h) (Conion et al., 2010)
	WTMM (Conlon et al., 2010)
	Decay index (Zuccarelio et al. 2014)
	Magnetic polarity inversion line characteristics (Mason & Hoeksema 2010)
	3D magnetic null point (Reid et al. 2012)
	R (Schrijver 2007) *
	LWL _{SG} (Falconer et al. 2008) *
	Ising energy (Ahmed et al. 2010)
	WG _M and S _H (Korsos et al. 2015)
	Magnetic helicity injection rate proxy (Park et al. 2010)

Spaceweather HMI Active Region Patches

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• From HMI full-disk Vector Magnetogram (12 min)

SD0/HMI Tracked AR (HARP) 2013/01/13 00:48

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SHARP vector components, LOS, continuum available at: http://jsoc.stanford.edu/ajax/lookdata.html

For this study: random selection of 25% of days.



Active region properties

- Active region properties considered:
 - <u>Schrijver's R value</u> (Flux PILs)
 - Fourier spectral index, α (Scaling of spectral power)
 - <u>MPIL-related properties</u> (Total length of PILs)
 - Decay index (Vert. variation of potential field)
 - Effective connected Field strength (coronal connectivity)
 - Ising energy (coronal connectivity)

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• Special emphasis on differences when using B_{los} and B_r .

Preliminary results

At 6 h cadence: 12733 regions. Valid points: 9454



Preliminary results

• Log(R)





Variation with AR longitudinal position

• Beff



• Alpha



Preliminary results

• Total length of MPIL



Summary

- FLARECAST is half way through and expecting to start producing forecasts by early 2017.
- About 70% of PE codes have been developed and tested. Processing of the SHARP DB will start soon.
- Coronal connectivity and fractal properties seem to be varies more with the B_z choice. MPIL-related properties are less sensitive.
- Four out six properties do not show strong dependence with AR longitude.
- Usefulness of properties will be determined by their flaring association.